

A NEW POINT OF VIEW ON GENERAL KALUZA-KLEIN THEORIES

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1. HISTORICAL COMMENTS AND MOTIVATION

- **MAXWELL** (1861): Electromagnetism = Electricity + magnetism.
- **GLASHOW, WEINBERG, SALAM** (1961): Electromagnetism + Weak Interactions

GRAVITY?

- **NORDSTRÖM** (1914): Electromagnetism + His Own Gravity Theory.
- **KALUZA** (1919): $5D$ Gravity Theory \implies $4D$ Gravity Theory and Electromagnetism (1921), "CYLINDER CONDITION"
- **KLEIN** (1926): "COMPACTIFICATION CONDITION" fifth dimension should be compact \implies KALUZA-KLEIN THEORY (KK THEORY).
- (1970): **STRING THEORY AND SUPERGRAVITY.**
- **EINSTEIN-BERGMANN** (1938).
- **WESSON** (1990) **STM THEORY.**
- **RANDALL-SUNDRUM** (1999) **BRANE WORLD THEORY.**

MOTIVATION: A New Approach on General KK Theories. (Remove Both Conditions; Use Nonholonomic Geometry).

2. GENERAL KALUZA-KLEIN SPACE

$$\bar{M} = M \times K, \quad \pi : \bar{M} \longrightarrow M$$

$$\tilde{x}^\alpha = \tilde{x}^\alpha(x^\mu), \quad \tilde{x}^4 = x^4 + h(x^\mu), \quad \alpha, \mu \in \{0, 1, 2, 3\} \quad (1)$$

\bar{g} - pseudo-Riemannian metric on \bar{M} , given locally by

$$g_{\alpha\beta}(x^i)dx^\alpha dx^\beta + \epsilon\Phi^2(x^i)(dx^4 + A_\alpha dx^\alpha)^2. \quad (2)$$

$V\bar{M}$ - vertical distribution spanned by $\partial/\partial x^4$.

$H\bar{M}$ - horizontal distribution spanned by

$$\frac{\delta}{\delta x^\alpha} = \frac{\partial}{\partial x^\alpha} - A_\alpha \frac{\partial}{\partial x^4}. \quad (3)$$

$$T\bar{M} = H\bar{M} \oplus V\bar{M}.$$

$$H\bar{M} \perp V\bar{M}$$

A_α - electromagnetic potentials

$g_{\alpha\beta}(x^i)$ - define a Lorentz metric g on $H\bar{M}$.

(\bar{M}, \bar{g}) = **general Kaluza-Klein space**.

(2) is invariant with respect to (1) if and only if:

$$g_{\alpha\beta} = \tilde{g}_{\mu\nu} \frac{\partial \tilde{x}^\mu}{\partial x^\alpha} \frac{\partial \tilde{x}^\nu}{\partial x^\beta}, \quad \text{and} \quad A_\alpha = \tilde{A}_\gamma \frac{\partial \tilde{x}^\gamma}{\partial x^\alpha} + \frac{\partial h}{\partial x^\alpha}.$$

3. 4D TENSOR FIELDS ON (\bar{M}, \bar{g})

$T = \left(T_{\beta_1 \dots \beta_q}^{\alpha_1 \dots \alpha_p}(x^i) \right)$ is a 4D tensor field of type (p, q) if:

$$T_{\beta_1 \dots \beta_q}^{\alpha_1 \dots \alpha_p} \frac{\partial \tilde{x}^{\gamma_1}}{\partial x^{\alpha_1}} \dots \frac{\partial \tilde{x}^{\gamma_p}}{\partial x^{\alpha_p}} = \tilde{T}_{\mu_1 \dots \mu_q}^{\gamma_1 \dots \gamma_p} \frac{\partial \tilde{x}^{\mu_1}}{\partial x^{\beta_1}} \dots \frac{\partial \tilde{x}^{\mu_q}}{\partial x^{\beta_q}}.$$

Examples: $g_{\alpha\beta}$, $g^{\alpha\beta}$.

- 4D electromagnetic tensor field

$$F_{\alpha\beta} = \frac{\delta A_\beta}{\delta x^\alpha} - \frac{\delta A_\alpha}{\delta x^\beta}; \quad \left[\frac{\delta}{\delta x^\beta}, \frac{\delta}{\delta x^\alpha} \right] = F_{\alpha\beta} \frac{\partial}{\partial x^4}.$$

- 4D Obstruction Tensor Fields:

$$D_{\alpha\beta} = \frac{1}{2} \frac{\partial g_{\alpha\beta}}{\partial x^4}, \quad B_\alpha = \frac{\partial A_\alpha}{\partial x^4}, \quad C_\alpha = \Phi^{-1} \frac{\delta \Phi}{\delta x^\alpha}.$$

(\bar{M}, \bar{g}) is the classical KK space $\iff D = 0, B = 0, C = 0$.

4. A REMARKABLE LINEAR CONNECTION ON $H\bar{M}$

$$\nabla : \Gamma(T\bar{M}) \times \Gamma(H\bar{M}) \rightarrow \Gamma(H\bar{M}); \quad (X, hY) \mapsto \nabla_X hY$$

Torsion Tensor Field: $T : \Gamma(T\bar{M}) \times \Gamma(H\bar{M}) \rightarrow \Gamma(H\bar{M});$

$$T(X, hY) = \nabla_X hY - \nabla_{hY} hX - h[X, hY].$$

Theorem 1. *There exists a unique linear connection ∇ on $H\bar{M}$ satisfying:*

- (i) ∇ is a metric connection ($\nabla_X g = 0$)
- (ii) The torsion tensor field of ∇ is given by

$$T(hX, hY) = 0, \quad T(vX, hY) = D(vX, hY),$$

∇ = Riemannian horizontal connection

$$(a) \quad \nabla_{hX} hY = h\bar{\nabla}_{hX} hY, \quad (b) \quad \nabla_{vX} hY = h[vX, hY] + D(vX, hY).$$

Theorem 2. *The Levi-Civita connection $\bar{\nabla}$ on (\bar{M}, \bar{g}) is expressed as follows:*

$$\bar{\nabla}_{\frac{\delta}{\delta x^\beta}} \frac{\delta}{\delta x^\alpha} = \Gamma_\alpha{}^\gamma{}_\beta \frac{\delta}{\delta x^\gamma} + \left(\frac{1}{2} F_{\alpha\beta} - \epsilon \Phi^{-2} D_{\alpha\beta} \right) \frac{\partial}{\partial x^4},$$

$$\bar{\nabla}_{\frac{\partial}{\partial x^4}} \frac{\delta}{\delta x^\alpha} = \left(D_\alpha{}^\gamma + \frac{\epsilon \Phi^2}{2} F_\alpha{}^\gamma \right) \frac{\delta}{\delta x^\gamma} + (C_\alpha - B_\alpha) \frac{\partial}{\partial x^4},$$

$$\bar{\nabla}_{\frac{\delta}{\delta x^\alpha}} \frac{\partial}{\partial x^4} = \left(D_\alpha{}^\gamma + \frac{\epsilon \Phi^2}{2} F_\alpha{}^\gamma \right) \frac{\delta}{\delta x^\gamma} + C_\alpha \frac{\partial}{\partial x^4},$$

$$\bar{\nabla}_{\frac{\partial}{\partial x^4}} \frac{\partial}{\partial x^4} = \epsilon \Phi^2 (B^\gamma - C^\gamma) \frac{\delta}{\delta x^\gamma} + \Psi \frac{\partial}{\partial x^4},$$

$$\Gamma_\alpha{}^\gamma{}_\beta = \frac{1}{2} g^{\gamma\mu} \left\{ \frac{\delta g_{\mu\alpha}}{\delta x^\beta} + \frac{\delta g_{\mu\beta}}{\delta x^\alpha} - \frac{\delta g_{\alpha\beta}}{\delta x^\mu} \right\}, \quad \Psi = \frac{1}{\Phi} \frac{\partial \Phi}{\partial x^4}.$$

5. 4D EQUATIONS OF MOTION IN (\bar{M}, \bar{g})

$\{\delta/\delta x^\alpha, \partial/\partial x^4\}$ - adapted frame field

Theorem 3. *The fully general equations of motion for the STM Theory on (\bar{M}, \bar{g}) are expressed as follows:*

$$\begin{aligned}
 (a) \quad & \frac{d^2x^\gamma}{dt^2} + \Gamma_\alpha{}^\gamma{}_\beta \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = \frac{\delta x^4}{\delta t} \left\{ \epsilon \Phi^2 (C^\gamma - B^\gamma) \frac{\delta x^4}{\delta t} \right. \\
 & \quad \left. - (2D_\alpha{}^\gamma + \epsilon \Phi^2 F_\alpha{}^\gamma) \frac{dx^\alpha}{dt} \right\}, \\
 (b) \quad & \frac{d}{dt} \left(\frac{\delta x^4}{\delta t} \right) - \epsilon \Phi^{-2} D_{\alpha\beta} \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} \\
 & = \frac{\delta x^4}{\delta t} \left\{ \Psi \frac{\delta x^4}{\delta t} + (B_\alpha - 2C_\alpha) \frac{dx^\alpha}{dt} \right\}.
 \end{aligned} \tag{5}$$

$\bar{C} \subset \bar{M}$, satisfying (5) is called *horizontal motion* if

$$\frac{\delta x^4}{\delta t} = \frac{dx^4}{dt} + A_\alpha \frac{dx^\alpha}{dt} = 0. \tag{6}$$

Corollary 1. \bar{C} is horizontal motion, if and only if, (6) and

$$(a) \quad \frac{d^2x^\gamma}{dt^2} + \Gamma_\alpha{}^\gamma{}_\beta \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = 0, \quad (b) \quad D_{\alpha\beta} \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = 0. \tag{7}$$

Corollary 2. Any horizontal motion is autoparallel of the Riemannian horizontal connection.

$C \subset M$, *h-motion* if it is projection of a horizontal motion $\bar{C} \subset \bar{M}$

Corollary 3. *Any motion from general relativity of the 4D spacetime (M, g) is the projection of some horizontal motions in (\bar{M}, \bar{g}) .*

$\bar{C} \subset \bar{M}$ is a *c-motion* if it satisfies

$$\frac{dx^4}{dt} + A_\alpha \frac{dx^\alpha}{dt} = c, \quad c \neq 0. \quad (8)$$

Corollary 4 *Any c-motion, in (\bar{M}, \bar{g}) is a solution of (8) and of the system*

$$(a) \quad \frac{d^2x^\gamma}{dt^2} + \Gamma_\alpha{}^\gamma{}_\beta \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = c \left\{ \epsilon c \Phi^2 (C^\gamma - B^\gamma) - (2D_\alpha{}^\gamma + \epsilon \Phi^2 F_\alpha{}^\gamma) \frac{dx^\alpha}{dt} \right\}, \quad (9)$$

$$(b) \quad D_{\alpha\beta} \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = \epsilon c \Phi^2 \left\{ (2C_\alpha - B_\alpha) \frac{dx^\alpha}{dt} - c\Psi \right\}.$$

Particular Case: Classical KK space: $D_{\alpha\beta} = 0$, $B_\alpha = 0$, $\Phi = const$.

Then take $c = -q/m$, q = charge, m = mass of particle.
Then (9) reduces to *Lorentz force equations*

$$\frac{d^2x^\gamma}{dt^2} + \Gamma_\alpha{}^\gamma{}_\beta \frac{dx^\alpha}{dt} \frac{dx^\beta}{dt} = \frac{q}{m} F_\alpha{}^\gamma \frac{dx^\alpha}{dt}.$$

Remark. Projections of $-q/m$ -motions in \bar{M} on M , bring more information than the solutions of Lorentz force equations.

6. FIFTH FORCE FROM FIFTH DIMENSION

\bar{C} geodesic in (\bar{M}, \bar{g}) given by: $x^i = x^i(t)$, $t \in [a, b]$.

$$\frac{d}{dt} = \frac{dx^\alpha}{dt} \frac{\delta}{\delta x^\alpha} + \frac{\delta x^4}{\delta t} \frac{\partial}{\partial x^4}.$$

Then we call $U(t) = \frac{dx^\alpha}{dt} \frac{\delta}{\delta x^\alpha}$ the *4D velocity* for *STM Theory*.

Case 1. \bar{C} is geodesic with spacelike or timelike *4D velocity*.

$s = 4D$ arc length parameter, that is, $ds^2 = \epsilon^* g_{\alpha\beta} dx^\alpha dx^\beta$. Then define the *fifth force* $\mathcal{F}(s)$ along \bar{C} by

$$\mathcal{F}(s) = \nabla_{\frac{d}{ds}} U(s).$$

∇ = Riemannian horizontal connection on $H\bar{M}$.

Similarly, the dual fifth force $\mathcal{F}^*(s)$ is given by

$$\mathcal{F}^*(s) = \nabla_{\frac{d}{ds}} U^*(s).$$

Properties: 1. $\mathcal{F}(s) \perp U(s)$, 2. $\mathcal{F}(s), \mathcal{F}^*(s)$ are indeed dual to each other.

Theorem 4. Let \bar{C} be horizontal geodesic of (\bar{M}, \bar{g}) .

Then we have:

- (i) The fifth force vanishes identically on (\bar{M}, \bar{g}) .
- (ii) The 4D arc length parameter is an affine parameter on \bar{C} .

Case 2. \bar{C} is geodesic in (\bar{M}, \bar{g}) with lightlike 4D velocity

$$\mathcal{F}(t) = \nabla_{\frac{d}{dt}} U(t).$$

Theorem 5. \bar{C} can be only in the following cases:

- (i) It is horizontal and null geodesic in (\bar{M}, \bar{g}) .
- (ii) It is non-null geodesic given by (5a) and the equation

$$\frac{d}{dt} \left(\frac{\delta x^4}{\delta t} \right) = \frac{\delta x^4}{\delta t} \left\{ \Psi \frac{\delta x^4}{\delta t} - C_\alpha \frac{dx^\alpha}{dt} \right\}.$$

7. RELATIVISTIC GENERAL KALUZA-KLEIN SPACE

(\bar{M}, \bar{g}) satisfying conditions

$$(A) \sum_{(\alpha, \beta, \gamma)} \{F_{\alpha\beta} D_\gamma^\mu\} = 0, \quad (B) \sum_{(\alpha, \beta, \gamma)} \{F_{\alpha\beta} R_\nu^\mu{}_{\gamma 4}\} = 0.$$

Curvature tensor field R of the Riemannian horizontal connection ∇ :

$$R \left(\frac{\delta}{\delta x^\gamma}, \frac{\delta}{\delta x^\beta} \right) \frac{\delta}{\delta x^\alpha} = R_\alpha^\mu{}_{\beta\gamma} \frac{\delta}{\delta x^\mu},$$

$$R \left(\frac{\partial}{\partial x^4}, \frac{\delta}{\delta x^\beta} \right) \frac{\delta}{\delta x^\alpha} = R_\alpha^\mu{}_{\beta 4} \frac{\delta}{\delta x^\mu},$$

$$4D \text{ Ricci tensor: } Ric(hX, hY) = \sum_{\alpha=0}^3 \epsilon_\alpha R(E_\alpha, hX, E_\alpha, hY)$$

$$4D \text{ scalar curvature: } \mathbf{R} = \sum_{\alpha=0}^3 \epsilon_\alpha Ric(E_\alpha, E_\alpha).$$

$$R_{\alpha\beta} = Ric \left(\frac{\delta}{\delta x^\beta}, \frac{\delta}{\delta x^\alpha} \right), R_{\alpha\beta} = R_\alpha^\mu{}_{\beta\mu}, \quad \mathbf{R} = g^{\alpha\beta} R_{\alpha\beta}.$$

$$\text{Proved: } 2R_{\alpha|\mu}^\mu = \mathbf{R}_{|\alpha}, \quad \mathbf{R}_{|\alpha} = \frac{\delta \mathbf{R}}{\delta x^\alpha}.$$

$$4D \text{ Einstein gravitational tensor field: } G_{\alpha\beta} = R_{\alpha\beta} - \frac{\mathbf{R}}{2} g_{\alpha\beta}.$$

Theorem 6. $G_{\alpha\beta}$ is 4D symmetric tensor field whose divergence vanishes identically on \bar{M} .

4D electromagnetic energy-momentum tensor field:

$$E_{\alpha\beta} = \frac{1}{4} g_{\alpha\beta} F_{\mu\nu} F^{\mu\nu} - F_{\alpha\mu} F_\beta^\mu.$$

8. 4D EINSTEIN EQUATIONS IN (\bar{M}, \bar{g})

$$\begin{aligned}
& \bar{G}_{ij} = 0, \quad i, j \in \{0, 1, 2, 3, 4\} \iff \\
(a) \quad & G_{\alpha\beta} = -\frac{\epsilon\Phi^2}{2}E_{\alpha\beta} + \frac{1}{2}F_{\mu\nu}F^{\mu\nu}g_{\alpha\beta} - \frac{1}{2}\left\{D_{\alpha\mu}F_{\beta}^{\mu} + F_{\alpha\mu}D_{\beta}^{\mu}\right\} + \epsilon\Phi^{-2}\left\{D_{\alpha\beta|4} + D_{\alpha\beta}D_{\mu}^{\mu} - \Psi D_{\alpha\beta}\right. \\
& \left.- \frac{1}{2}g_{\alpha\beta}\left((D_{\mu}^{\mu})^2 - D_{\mu\nu}D^{\mu\nu}\right)\right\} + \frac{1}{2}\left\{C_{\alpha|\beta} + C_{\beta|\alpha}\right. \\
& \left.- B_{\alpha|\beta} - B_{\beta|\alpha}\right\} + (B_{\alpha} - C_{\alpha})(B_{\beta} - C_{\beta}), \\
(b) \quad & \frac{\epsilon\Phi^2}{2}\left\{F_{\alpha}^{\mu}|_{\mu} + 3F_{\alpha}^{\mu}C_{\mu} - 2F_{\alpha}^{\mu}B_{\mu}\right\} = D_{\mu}^{\mu}|_{\alpha} \quad (10) \\
& - D_{\alpha}^{\mu}|_{\mu} + D_{\alpha}^{\mu}C_{\mu} - D_{\mu}^{\mu}C_{\alpha}, \\
(c) \quad & \frac{\epsilon\Phi^2}{4}F_{\mu\nu}F^{\mu\nu} + \epsilon\Phi^{-2}\left\{\Psi D_{\mu}^{\mu} - D_{\mu\nu}D^{\mu\nu}\right. \\
& \left.- D_{\mu}^{\mu}|_4\right\} + B^{\mu}|_{\mu} - C^{\mu}|_{\mu} \\
& -(B^{\mu} - C^{\mu})(B_{\mu} - C_{\mu}) = 0.
\end{aligned}$$

(10a) represents the 4D Einstein equations in (\bar{M}, \bar{g}) .

Wesson-Ponce de Leon Theory ($A_{\alpha} = 0$).

$$\begin{aligned}
G_{\alpha\beta} = & \Phi^{-1}\Phi_{\alpha|\beta} - \frac{\epsilon}{2}\Phi^{-2}\left\{\Phi^{-1}\overset{*}{\Phi}\overset{*}{g}_{\alpha\beta} - \overset{**}{g}_{\alpha\beta} + \overset{*}{g}^{\mu\nu}g_{\mu\alpha}\overset{*}{g}_{\nu\beta}\right. \\
& \left.- \frac{1}{2}\overset{*}{g}_{\alpha\beta}g^{\mu\nu}\overset{*}{g}_{\mu\nu} + \frac{1}{4}g_{\alpha\beta}\left(\overset{*}{g}^{\mu\nu}\overset{*}{g}_{\mu\nu} + \left(g^{\mu\nu}\overset{*}{g}_{\mu\nu}\right)^2\right)\right\}.
\end{aligned}$$

\star means usual partial derivative with respect to x^4 .

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