

# **Systems with a Position-Dependent-Mass and Symmetry-Preserving Inverse Problems in Lagrangian Dynamics**

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# RÉSUMÉ

- ▶ CLASSICAL INVERSE PROBLEMS
- ▶ INVERSE PROBLEM INVOLVING NOETHER-SYMMETRY
- ▶ CONSTANTS OF THE MOTION OF PERTURBATIONS
- ▶ A PDM-VERSION OF ERMAKOV-RAY-REID SYSTEMS
- ▶ A FUNCTIONAL APPROACH
- ▶ PERTURBED PDM-SYSTEMS IN  $N = 3$
- ▶ PERTURBATIONS ASSOCIATED TO FLAT METRICS
- ▶ PDM-SYSTEMS WITH DISSIPATIVE TERMS
- ▶ APPROACH VIA REALIZATIONS OF VF
- ▶ AN  $n$ -DIMENSIONAL EXTRAPOLATION
- ▶ SOME FINAL REMARKS. POSSIBLE OUTLINES.

## CLASSICAL INVERSE PROBLEMS

- (Flat) Bertrand theorem: If  $\mathbf{F} = -\frac{dU}{dt}$  central and “bounded orbit” implies “closed orbit”  $\Rightarrow U = kr^{-1}$  or  $U = kr^2$
- First example of large hierarchy of different inverse problems.<sup>1</sup>

<sup>1</sup>Bertrand J 1852 *J. Math. Pures Appl.* **17** 121; Bertrand J 1877 *Comp. Rend.* **84** 671; Dainelli U 1880 *Giorn. Mat.* **18** 271; Imshenetsky V G 1882 *Mém. Soc. Sci. Bordeaux* **4** 31; von Helmholtz H 1887 *J. reine angew. Math.* **100** 137; Suslov G K. On a force function admitting given integrals. Kiev 1890; Ermakov V P 1891 *Recueil Math.* **15** 611; Darboux G 1901 *Arch. Neerl.* **6** 371; Dall'Acqua F A 1908 *Math. Annalen* **66** 398; Kasner E 1909 *Differential Geometric Aspects of Dynamics* (AMS, New York); Eisenhart LP 1909 *A Treatise on the Differential Geometry of Curves and Surfaces* (Dover, New York); Burgatti P 1911 *Rend. Acad. Lincei* **20** 108; Schaefer C 1919 *Die Prinzipie de Dynamik* (W. de Gruyter, Berlin); Charlier C L 1927 *Die Mechanik des Himmels* (Veit, Leipzig); Douglas D 1941 *Trans. Amer. Math. Soc.* **50** 71; Whittaker E T 1944 *A Treatise on the Analytical Dynamics of Particles and Rigid Bodies* (Dover Pub., New York); Chaplygin S A. Sobr. soch. t. 1, GITTL 1948; Erugin N P 1952 *PMM* **16** 59; Drâmbă C 1963 *St. Cerc. Astron. Acad. Rep. Pop. Rom* **8** 7; Steudel H 1967 *Ann. Physik* **475** 11; Szebehely V G 1967 *Theory of Orbits* (Academic Press, New York); Galiullin A S 1972 *Differenc. uravneniya* **8** 1535; Djukić Dj S, Vujanović B D 1975 *Acta Mech.* **23** 17; Rosenberg R 1977 *Analytical Dynamics of Discrete Systems* (Plenum Press, New York); Santilli R M 1978 *Foundations of Theoretical Mechanics vol. I* (Springer, New York); Szebehely V, Broucke R 1981 *Celestial Mechanics* **30** 395; Галиуллин А С 1981 *Обратные задачи динамики*, Наука, Москва; Mertens R 1981 *ZAMP* **61** T252; Kolokol'tsev V 1983 *Izv. Akad. Nauk SSSR* **46** 994; Stavinschi M 1985 *Topics Astrophys. Astron. Sp. Sci.* **1** 41; Melis A, Borghero F 1986 *Meccanica* **21** 71; Bozis G, Ichtiaoglou S 1987 *Inverse Problems* **3** 213; Shorokhov S G 1988 *Celestial Mechanics* **44** 193; Kozlov VV, Harin A O 1992 *Celestial Mechanics* **54** 393; Bozis G 1995 *Inverse Problems* **11** 687; Mei F X 2000 *Acta Mech.* **141** 135; Ziglin S L 2001 *Doklady Phys.* **46** 570; Borisov AV, Mamaev IS 2004 *Reg. Chaotic Dyn.* **9** 265; Voyatzsi D, Ichtiaoglou S 2005 *Celestial Mechanics* **93** 331; Kotoulas T, Ichtiaoglou S 2006 *J. Geom. Phys.* **56** 2447; Mušicki, Dj. 2012 *Acta Mech.* **223** 2117; RCS 2014 *Comm. Nonlinear Sci. Num. Simulat.* **19** 2602; Casetta L, Pesce C P 2015 *Acta Mech.* **226** 63; RCS 2016 *Acta Mech.* **227** 1941.

## Dainelli and Zhukovski problems

- Given plane curves  $\varphi(x, y) = \alpha$ , find the most general forces such that  $\varphi$  is an admissible trajectory.
- Components:<sup>2</sup>

$$F_x = u(x, y) (\varphi_{xy}\varphi_y - \varphi_{yy}\varphi_x) + \frac{1}{2}\varphi_y \left( \varphi_y \frac{\partial u}{\partial x} - \varphi_x \frac{\partial u}{\partial y} \right), \quad (1)$$

$$F_y = u(x, y) (\varphi_{xy}\varphi_x - \varphi_{xx}\varphi_y) + \frac{1}{2}\varphi_x \left( \varphi_x \frac{\partial u}{\partial y} - \varphi_y \frac{\partial u}{\partial x} \right). \quad (2)$$

- No constraint on  $u(x, y)$ . When  $\mathbf{F} = -\frac{\partial V}{\partial x} \frac{\partial}{\partial x} - \frac{\partial V}{\partial y} \frac{\partial}{\partial y}$  ?
- Explicitly:

$$\begin{aligned} & \frac{1}{2}\varphi_x\varphi_y (u_{xx} - u_{yy}) + \varphi_y \left( u\varphi_{xy}^2 + u_x\varphi_{y^2} + \frac{3}{2}u_x\varphi_{x^2} + \frac{1}{2}u_y\varphi_{xy} + u\varphi_{x^3} \right) \\ & - \varphi_x \left( u\varphi_{y^3} + \frac{1}{2}u_x\varphi_{xy} + \frac{3}{2}u_y\varphi_{y^2} + u_y\varphi_{x^2} + u\varphi_{yx^2} \right) \frac{1}{2}u_{xy} (\varphi_y^2 - \varphi_x^2) = 0. \end{aligned} \quad (3)$$

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<sup>2</sup>Normalized units ( $m = 1$ ) and  $u(x, y) = k(x, y)^2$ .

- Extrapolation to manifolds (surfaces in  $\mathbb{R}^3$ ) due to Zhukovsky
- Free motion on surface  $S$  with line element  $\overline{ds^2}$

$$L = T = \frac{1}{2} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) = \frac{1}{2} (E \dot{u}^2 + 2F \dot{u} \dot{v} + G \dot{v}^2) \quad (4)$$

- Type II Lagrangian equations:

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}^i} \right) - \left( \frac{\partial L}{\partial q^i} \right) = 0, \quad i = 1, 2; \quad (q^1, q^2) = (u, v) \quad (5)$$

- Using the associated linear connection

$$\begin{aligned} \Gamma_{11}^1 &= \frac{1}{2\Delta} \left( G \frac{\partial E}{\partial u} + F \frac{\partial E}{\partial v} - 2F \frac{\partial F}{\partial u} \right), & \Gamma_{12}^1 &= \frac{1}{2\Delta} \left( G \frac{\partial E}{\partial v} - F \frac{\partial G}{\partial u} \right), \\ \Gamma_{22}^1 &= \frac{1}{2\Delta} \left( 2G \frac{\partial F}{\partial v} - G \frac{\partial G}{\partial u} - F \frac{\partial G}{\partial v} \right), & \Gamma_{11}^2 &= \frac{1}{2\Delta} \left( 2E \frac{\partial F}{\partial u} - F \frac{\partial E}{\partial u} - E \frac{\partial E}{\partial v} \right), \\ \Gamma_{12}^2 &= \frac{1}{2\Delta} \left( E \frac{\partial G}{\partial u} - G \frac{\partial E}{\partial v} \right), & \Gamma_{22}^2 &= \frac{1}{2\Delta} \left( F \frac{\partial G}{\partial u} + E \frac{\partial G}{\partial v} - 2F \frac{\partial F}{\partial v} \right). \end{aligned} \quad (6)$$

► geodesic equations:

$$\ddot{u} + \Gamma_{11}^1 \dot{u}^2 + 2\Gamma_{12}^1 \dot{u} \dot{v} + \Gamma_{22}^1 \dot{v}^2 = 0, \quad (7)$$

$$\ddot{v} + \Gamma_{11}^2 \dot{u}^2 + 2\Gamma_{12}^2 \dot{u} \dot{v} + \Gamma_{22}^2 \dot{v}^2 = 0. \quad (8)$$

► Given a regular surface  $S = S(u, v)$  with line element  $\overline{ds^2}$ , the most general potential such that  $u = \text{cte}$  is an admissible trajectory for the holonomic constraint of moving on  $S$  has the form

$$U(u, v) = \frac{\zeta(v)}{G(u, v)} + \frac{1}{G(u, v)} \int \varphi(u) \frac{\partial G}{\partial u} du. \quad (9)$$

- Arbitrary functions  $\zeta(v)$  y  $\varphi(u)$ .
- $G^{-1}$  can be identified with the first Beltrami parameter.

## The Szebehely-Mertens equation

► Motivation: Determine potentials for which an observed trajectory  $\varphi(u, v) = C$  is contained in a given surface  $S$ . <sup>3</sup>

►  $u, v$  not necessarily orthogonal:  $g = \overline{ds^2} = Edu^2 + 2Fdudv + Gdv^2$

► From  $\frac{d}{dt}\varphi(u, v) = \varphi_u\dot{u} + \varphi_v\dot{v} = 0$  we get

$$\dot{u} = H(u, v)(F\varphi_u + G\varphi_v), \quad \dot{v} = -H(u, v)(E\varphi_u + F\varphi_v) \quad (10)$$

► Components

$$\begin{aligned} \ddot{u} &= (H_u\dot{u} + H_v\dot{v})(F\varphi_u + G\varphi_v) + H(u, v)(\varphi_u(F_u\dot{u} + F_v\dot{v}) + F(\varphi_u^2\dot{u} + \varphi_{uv}\dot{v})) + \\ &\quad H(u, v)(\varphi_v(G_u\dot{u} + G_v\dot{v}) + G(\varphi_{uv}\dot{u} + \varphi_{v^2}\dot{v})) \end{aligned} \quad (11)$$

$$\begin{aligned} -\ddot{v} &= (H_u\dot{u} + H_v\dot{v})(E\varphi_u + F\varphi_v) + H(u, v)(\varphi_u(E_u\dot{u} + E_v\dot{v}) + E(\varphi_u^2\dot{u} + \varphi_{uv}\dot{v})) + \\ &\quad H(u, v)(\varphi_v(F_u\dot{u} + F_v\dot{v}) + F(\varphi_{uv}\dot{u} + \varphi_{v^2}\dot{v})) \end{aligned} \quad (12)$$

► Assumption: There exist  $E_0(\varphi(u, v))$  y  $U(u, v)$  satisfying

$$U(u, v) = \frac{1}{2}g(\mathbf{w}, \mathbf{w}) - E_0(\varphi(u, v)); \quad \mathbf{w} = (\dot{u}, \dot{v}) \quad (13)$$

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<sup>3</sup>Considered by Szebehely for the planar case (angular momentum conservation in satellite systems), generalization due, among others, to Mertens R 1981 ZAMM **61** T252. See also Shorokhov S G 1998 *Celestial Mechanics* **44** 193.

► As  $E_0 = \text{cons}$  on each orbit  $\varphi(u, v) = \alpha \implies \frac{\partial E_0}{\partial u} = \frac{\partial E_0}{\partial \varphi} \frac{\partial \varphi}{\partial u}$  y  $\frac{\partial E_0}{\partial v} = \frac{\partial E_0}{\partial \varphi} \frac{\partial \varphi}{\partial v}$ .

► Now, from (13) we deduce that

$$\frac{\partial U}{\partial u} = \frac{1}{2} \frac{\partial}{\partial u} (g(\mathbf{w}, \mathbf{w})) - \frac{\partial E_0}{\partial \varphi} \frac{\partial \varphi}{\partial u} = 0; \quad \frac{\partial U}{\partial v} = \frac{1}{2} \frac{\partial}{\partial v} (g(\mathbf{w}, \mathbf{w})) - \frac{\partial E_0}{\partial \varphi} \frac{\partial \varphi}{\partial v} = 0.$$

► Developing this expression, taking into account (10), (11) y (12)<sup>4</sup>

$$\left( G \frac{\partial \varphi}{\partial u} - F \frac{\partial \varphi}{\partial v} \right) \frac{\partial U}{\partial u} + \left( E \frac{\partial \varphi}{\partial v} - F \frac{\partial \varphi}{\partial u} \right) \frac{\partial U}{\partial v} = 2W (E_0 - U) \quad (14)$$

► where

$$W = \frac{1}{A} \left( \Delta (\varphi_v^2 \varphi_{u^2} - 2\varphi_u \varphi_v \varphi_{uv} + \varphi_u^2 \varphi_{v^2}) - B_1 (G\varphi_u - F\varphi_v) - B_2 (E\varphi_v - F\varphi_u) \right)$$

► and  $A = E\varphi_v^2 - 2F\varphi_u \varphi_v + G\varphi_u^2$ ,  $B_1 = \frac{1}{2}E_u \varphi_v^2 - E_v \varphi_u \varphi_v + (F_v - \frac{1}{2}G_u) \varphi_u^2$ ,

►  $B_2 = \frac{1}{2}G_v \varphi_u^2 - G_u \varphi_u \varphi_v + (F_u - \frac{1}{2}E_v) \varphi_v^2$ .

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<sup>4</sup>Paraphrasing Eliè Cartan: “Il se trouve que c'est un calcul extrêmement pénible!”.

SPECIAL CASE:  $S = \mathbb{R}^2$   $E = G = 1$ ,  $F = 0$ ,  $u = x$ ,  $v = y$ : From (15)

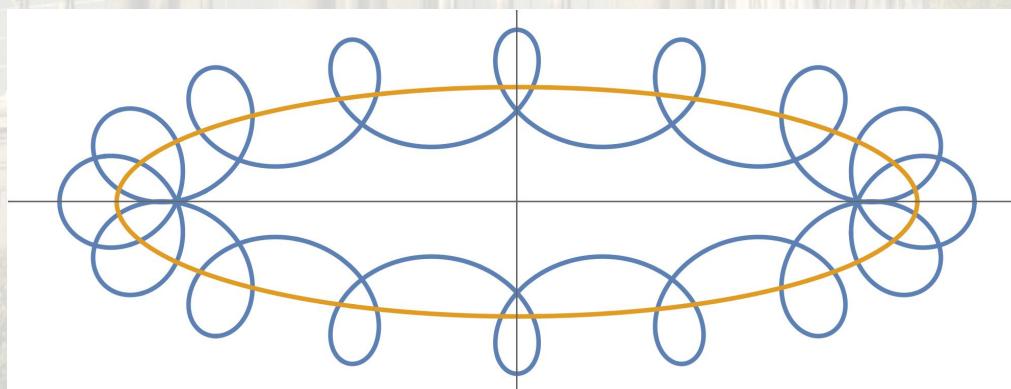
- $\Delta = 1$ ,  $A = \varphi_u^2 + \varphi_v^2$ ,  $B_1 = B_2 = 0$ ,

$$W = \frac{\varphi_{xx}\varphi_{x^2} - 2\varphi_{xy}\varphi_x\varphi_y + \varphi_{y^2}\varphi_x^2}{\varphi_x^2 + \varphi_y^2}$$

- and we get the “classical” Szebehely equation:<sup>5</sup>

$$\frac{\partial \varphi}{\partial x} \frac{\partial U}{\partial x} + \frac{\partial \varphi}{\partial y} \frac{\partial U}{\partial y} = 2 \frac{\varphi_{xx}\varphi_{x^2} - 2\varphi_{xy}\varphi_x\varphi_y + \varphi_{y^2}\varphi_x^2}{\varphi_x^2 + \varphi_y^2} (E_0 - U(x, y)) \quad (15)$$

- prototype orbit to be analyzed




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<sup>5</sup>Szebehely V G. *On the determination of the potential by satellite observations*. In: Proverbio E (Ed.) Proceedings Internat. Meeting on “Earth’s Rotation by Satellite Observation”, Bologna 1973, pp. 31-35.

**Solution without energy prescription**  $E_0 = E_0(\varphi(u, v))$

- Use the auxiliary function

$$\alpha = \frac{1}{2W} \left( G \frac{\partial \varphi}{\partial u} - F \frac{\partial \varphi}{\partial v} \right); \quad \beta = \left( E \frac{\partial \varphi}{\partial v} - F \frac{\partial \varphi}{\partial u} \right); \quad \gamma = \frac{\partial \varphi}{\partial v} \left( \frac{\partial \varphi}{\partial u} \right)^{-1} \quad (16)$$

- Energy

$$E_0 = V + \alpha \frac{\partial U}{\partial u} + \beta \frac{\partial U}{\partial v}. \quad (17)$$

- Relations:

$$\frac{\partial E_0}{\partial u} = \frac{\partial E_0}{\partial \varphi} \frac{\partial \varphi}{\partial u}; \quad \frac{\partial E_0}{\partial v} = \frac{\partial E_0}{\partial \varphi} \frac{\partial \varphi}{\partial v} = \gamma \frac{\partial E_0}{\partial u}.^6 \quad (18)$$

- The last identity and (17) lead to the hyperbolic equation

$$\alpha \gamma \frac{\partial^2 U}{\partial u^2} + (\beta \gamma - \alpha) \frac{\partial^2 U}{\partial u \partial v} - \beta \frac{\partial^2 U}{\partial v^2} + (\gamma + \gamma \alpha_u - \alpha_v) \frac{\partial U}{\partial u} + (\gamma \beta_u - \beta_v - 1) \frac{\partial U}{\partial v} = 0. \quad (19)$$

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<sup>6</sup>Here we assume that the condition  $\frac{\partial \varphi}{\partial u} \frac{\partial \varphi}{\partial v} \neq 0$  holds. For the degenerate case see [Bozis G, Mertens R 1985 ZAMM 65 383].

## The Drâmbă equation

- Describes motion in synodic references ( $\omega = 1$ ):<sup>7</sup>

$$\ddot{u} - 2\dot{v} = \Omega_u, \quad \ddot{v} + 2\dot{u} = \Omega_v, \quad (20)$$

- Potential  $\Omega$  includes the centrifugal term

$$\Omega = \frac{1}{2} (\dot{u}^2 + \dot{v}^2) - U(u, v) \Rightarrow \dot{u}^2 + \dot{v}^2 = 2\Omega - C \quad (21)$$

- In such a reference, (15) is given by<sup>8</sup>

$$\varphi_u \Omega_u + \varphi_v \Omega_v + W(2\Omega - C) \mp 2\sqrt{2\Omega - C} \sqrt{\varphi_u^2 + \varphi_v^2} = 0. \quad (22)$$

- “Disadvantage”: truly non-linear equation.

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<sup>7</sup>Follow at once from

$$m \ddot{\mathbf{r}}' = \mathbf{F} - m \ddot{\mathbf{r}}_0 - m \dot{\boldsymbol{\omega}} \times \mathbf{r}' - m \boldsymbol{\omega} \times (\boldsymbol{\omega} \times \mathbf{r}') - 2m \boldsymbol{\omega} \times \dot{\mathbf{r}}'.$$

with the values  $\mathbf{r}_0 = 0$ ,  $\boldsymbol{\omega} = \mathbf{e}_3$  and  $m = 1$ .

<sup>8</sup>Originally proposed by Drâmbă C 1963 *St. Cerc. Astron. Acad. Rep. Pop. Rom.* **8** 7. See also Stavinschi M, Mioc V 1993 *Astron. Nachr.* **314** 91.

## Generic orbit problem on pseudo-Riemannian manifolds

- $M$  manifold of dimension  $n$ ,  $g = g_{ij} dq^i \otimes dq^j$  nondegenerate metric tensor.
- $\mathfrak{X}(M)$  VF on  $M$ ,  $\Omega^1(M)$  1-forms,  $\nabla$  Levi-Civita connection.
- $X \in \mathfrak{X}(M)$  is an integral of  $\omega \in \Omega^1(M)$  if  $\omega(X) = 0$ .
- Given independent forms  $\{w_1, \dots, \omega_r\}$ ,<sup>9</sup> we consider

$$\mathcal{D}(X) = \{X \in \mathfrak{X}(M) \mid \omega_j(X) = 0, j = 1..r\} \quad (23)$$

- Extend to a (local) basis  $\{w_1, \dots, \omega_r, \theta_1, \dots, \theta_{n-r}\} \subset \Omega^1(M)$ , and take

$$\Xi = \omega_1 \wedge \omega_1 \wedge \dots \wedge \omega_r \wedge \theta_1 \wedge \dots \wedge \theta_{n-r}, \quad (24)$$

- Volume form  $\widehat{\Xi} = \Xi \left( \frac{\partial}{\partial q_1}, \frac{\partial}{\partial q_2}, \dots, \frac{\partial}{\partial q_n} \right) \neq 0$

- NOTATION:

$$\{w_1, \dots, \omega_r, \theta_1, \dots, \theta_{n-r}\} = \{\Lambda_1, \dots, \Lambda_n\} \quad (25)$$

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<sup>9</sup>Distribution of order  $n - r$ .

- generic form of annihilating VF  $X$ :  $(\frac{\partial}{\partial q^k} = \partial_{q^k})$

$$X = \det \begin{pmatrix} \omega_1(\partial_{q^1}) & \omega_1(\partial_{q^2}) & \cdots & \omega_1(\partial_{q^n}) & 0 \\ \omega_2(\partial_{q^1}) & \omega_2(\partial_{q^2}) & \cdots & \omega_2(\partial_{q^n}) & 0 \\ \vdots & \vdots & & \vdots & \vdots \\ \omega_r(\partial_{q^1}) & \omega_r(\partial_{q^2}) & \cdots & \omega_r(\partial_{q^n}) & 0 \\ \theta_1(\partial_{q^1}) & \theta_1(\partial_{q^2}) & \cdots & \theta_1(\partial_{q^n}) & X^{r+1} \\ \vdots & \vdots & & \vdots & \vdots \\ \theta_{n-r}(\partial_{q^1}) & \theta_{n-r}(\partial_{q^2}) & \cdots & \theta_{n-r}(\partial_{q^n}) & X^n \\ \partial_{q^1} & \partial_{q^2} & \cdots & \partial_{q^n} & 0 \end{pmatrix} \quad (26)$$

- Corresponding 1-form  $\omega_X$  associated to  $X$ :

$$\omega_X = g_{jk} X^k dq^k \quad (27)$$

- Coboundary operator

$$d\omega_X = \frac{1}{2} \left( \frac{\partial p_j}{\partial q^k} - \frac{\partial p_k}{\partial q^j} \right) dq^j \wedge dq^k = \frac{1}{2} a^{jk} \Lambda_j \wedge \Lambda_k \quad (28)$$

- $a^{jk}$  given in terms of  $\widehat{\Xi}$

$$a^{jk} = (-1)^{j+k-1} \widehat{\Xi}^{-1} d\omega_X \wedge \Lambda_1 \wedge \cdots \wedge \widehat{\Lambda}_k \wedge \cdots \wedge \cdots \wedge \widehat{\Lambda}_j \wedge \cdots \wedge \Lambda_n \left( \frac{\partial}{\partial q^1}, \dots, \frac{\partial}{\partial q^n} \right)$$

- Auxiliary condition:

$$X \rfloor d\omega_X \left( \frac{\partial}{\partial q^k} \right) = \Psi^j \Lambda_j \left( \frac{\partial}{\partial q^k} \right) \quad (29)$$

- System on  $M$  subjected to the constraint  $\omega = F_k(\mathbf{q}, \dot{\mathbf{q}}) dq^k$  with  $T = \frac{1}{2} g_{ij} \dot{q}^i \dot{q}^j$ :

$$\frac{d}{dt} \left( \frac{\partial T}{\partial \dot{q}^k} \right) - \left( \frac{\partial T}{\partial q^k} \right) = Q_k = \omega \left( \frac{\partial}{\partial q^k} \right), \quad (30)$$

- If  $\omega$  adopts the form

$$\omega = \frac{1}{2} d(g(X, X)) + \sum_{j=1}^r \Psi^j \Lambda_j \quad (31)$$

- Then the trajectories of the system are those determined by the first order system with constraints  $\Psi^j = 0$  for  $j = r+1, \dots, n$

$$\dot{\mathbf{q}} = X(\mathbf{q}) \quad (32)$$

- Formally, (30) with  $\omega$  of the type (31) corresponds to a non-holonomic system constrained by the Pfaff form

$$\sum_{l=1}^n \Lambda_j \left( \frac{\partial}{\partial q^l} \right) \dot{q}^l = 0, \quad j = 1, \dots, r. \quad (33)$$

- If  $\Lambda_j = d\varphi_j$  (Fröbenius), the system is holonomic.
- If  $r = n - 1$ , the most general force field giving rise to the orbits

$$\varphi_j = \alpha_j, \quad j = 1, \dots, r = n - 1. \quad (34)$$

is given by

$$\omega = \frac{1}{2} d(g(X, X)) + \rho(\mathbf{q}) \sum_{j=1}^{n-1} a_{nj} d\varphi_j \quad (35)$$

- The case (35) amplifies the methods of Szebehely, Mertens and Drâmbă.
- Applications: Nonholonomic systems.<sup>10</sup>

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<sup>10</sup>Klein J 1962 *Ann. Inst. Fourier* **12** 1; Selig J M 1996 *Geometric Fundamentals of Robotics* (Springer, New York); Fedorov Yu N, Jovanović B D 2004 *J. Nonlinear Sci.* **14** 341; Bloch A M 2005 *Nonholonomic Dynamics and Control Theory* (Springer, Berlin); Razavy M 2006 *Classical and Quantum Dissipative Systems* (Imperial College Press, London)

## THE ERUGIN APPROACH

- Further Inverse problems in terms of quantitative or qualitative restrictions in the configuration or phase space (and its generalization in the Cartan sense)
- Two possibilities basing on “integral” manifolds:  $\mathbf{q} = (q^1, \dots, q^n)$

$$\omega_\mu(t, \mathbf{q}, \dot{\mathbf{q}}) = \lambda_\mu; \quad 1 \leq \mu \leq n \quad (36)$$

- $\omega_\mu(t, \mathbf{q}, \dot{\mathbf{q}})$  compatible and independent.

1. Construct the equations of motion  $\ddot{q}_\nu = \Phi(t, \mathbf{q}, \dot{\mathbf{q}}), \quad 1 \leq \nu \leq n$  in accordance with (36). Analyze the existence conditions for (generalized) potentials.
2. Given a system  $\ddot{q}_\nu = \Phi(t, \mathbf{q}, \dot{\mathbf{q}}, \mathbf{v}), \quad 1 \leq \nu \leq n$ , determine the system parameters  $\mathbf{v}$  compatible with (36) [system restauration]

- Both (and additional) possibilities reducible to a formulation in terms of DE (original solution due to Erugin)<sup>11</sup>

<sup>11</sup>Erugin N P 1952 Prikl. Mat. Mekh. **16** 59; Galiullin A S 1960 *Some Questions about the Stability of Programmed Motion*, (KGU, Kazan); Dragović V, Gajić B, Jovanović B D 1998 *J. Phys. A: Math. Gen.* **31** 9861; Sadovskaya N, Ramírez R 2004 *J. Phys. A: Math. Gen.* **35** 3847; Ramírez R, Sadovskaya N 2007 *Rep. Math. Phys.* **60** 427; Borisov A V, Mamaev I S 2008 *Regul. Chaotic Dyn.* **13** 443; Soltakhanov Sh Kh, Yushkov M, Zegzhda S 2009 *Mechanics of Non-holonomic Systems. A New Class of Control Systems* (Springer, New York)

## THE FIRST NOETHER THEOREM<sup>12</sup>

► Given the action integral

$$W = \int_{t_0}^{t_1} L(t, \mathbf{q}, \dot{\mathbf{q}}) dt, \quad (37)$$

the total variation  $\Delta W$  of  $W$  with respect to an infinitesimal transformation

$$\bar{t} = t + \Delta t; \quad \bar{\mathbf{q}} = \mathbf{q} + \Delta \mathbf{q}$$

is given by

$$\Delta W = \int_{t_0}^{t_1} \sum_{k=1}^N \frac{\delta L}{\delta q_k} (\Delta q_k - \dot{q}_k \Delta t) dt + \int_{t_0}^{t_1} \frac{d}{dt} \left[ L \Delta t + \sum_{k=1}^N \left( \frac{\partial L}{\partial \dot{q}_k} \Delta q_k - \dot{q}_k \frac{\partial L}{\partial \dot{q}_k} \Delta t \right) \right] dt, \quad (38)$$

with

$$\frac{\delta L}{\delta q_k} = \left( \frac{\partial L}{\partial q_k} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}_k} \right)$$

the variational derivatives of  $L$  with respect to  $q_k$

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<sup>12</sup>Noether E 1918 *Nachr. Ges. Wiss. Göttingen, Math.-Phys. Kl.* **1918** 235; Bessel-Hagen E 1921 *Math. Annalen* **84** 258; Funk P 1962 *Variationsrechnung und ihre Anwendung in Physik und Technik* (Berlin: Springer)

- $r$ -parameter group  $\{\epsilon_1, \dots, \epsilon_r\}$
- Invariance of  $W$  up to a divergence:

$$\bar{t} = t + \sum_{\sigma=1}^r \epsilon_\sigma \xi^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}); \quad \bar{q}_k = q_k + \sum_{\sigma=1}^r \epsilon_\sigma \eta_k^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}), \quad k = 1, \dots, N \quad (39)$$

- There exist functions (gauge terms)  $V^\sigma(t, \mathbf{q}, \dot{\mathbf{q}})$  for  $1 \leq \sigma \leq r$  such that

$$\Delta W \left[ \sum_{\sigma=1}^r \epsilon_\sigma \xi^\sigma, \sum_{\sigma=1}^r \epsilon_\sigma \eta^\sigma \right] = \int_{t_0}^{t_1} \frac{d}{dt} \left( \sum_{\sigma=1}^r \epsilon_\sigma V^\sigma \right) \quad (40)$$

- Absolute invariance:  $V^\sigma = 0$  for all  $\sigma$ .
- The Noether theorem does **not** impose the symmetry generator to be a mere transformation of the (extended) configuration space, as has been “popularized” after Hill’s paper.<sup>13</sup>

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<sup>13</sup>Hill E L 1951 *Rev. Mod. Phys.* **23** 253; Komorowski J 1968 *Studia Math.* **29** 261; Khukhunashvili Z V 1968 *Soviet Phys.* **11** 1; Rund H 1972 *Utilitas Math.* **2** 205; Havas P 1973 *Acta Phys. Austriaca* **38** 145; Logan J D 1973 *Aequat. Math.* **9** 210; Logan J D 1974 *J. Math. Anal. Appl.* **42** 191; Logan J D, Blakeslee J S 1975 *J. Math. Phys.* **16** 1374; Kosmann-Schwarzbach Y 2011 *The Noether Theorems* (New York: Springer Verlag)

► **First Noether Theorem.** *If  $W$  is invariant up to a divergence under the  $r$ -parameter group of transformations (39), then  $r$  linearly independent combinations of the variational derivatives become divergences, i.e.,*

$$-\sum_{k=1}^N \frac{\delta L}{\delta q_k} (\eta_k^\sigma - \dot{q}_k \xi^\sigma) = \frac{d}{dt} \left[ \left( L - \sum_{k=1}^N \dot{q}_k \frac{\partial L}{\partial \dot{q}_k} \right) \xi^\sigma + \sum_{k=1}^N \frac{\partial L}{\partial \dot{q}_k} \eta_k^\sigma - V^\sigma \right], \quad (41)$$

where  $\sigma = 1, \dots, r$ .

► Converse also holds. If  $r$  linearly independent combinations of the variational derivatives are divergences,  $W$  is invariant with respect to an  $r$ -parameter group of transformations.

► Variation problem:  $\frac{\delta L}{\delta q_k} = 0$  ( $1 \leq k \leq N$ ) implies that

$$\Psi^\sigma \doteq \left[ L - \sum_{k=1}^N \dot{q}_k \frac{\partial L}{\partial \dot{q}_k} \right] \xi^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}) + \sum_{k=1}^N \frac{\partial L}{\partial \dot{q}_k} \eta_k^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}) - V^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}), \quad \sigma = 1, \dots, r \quad (42)$$

are constants of the motion.

► Expanding  $\frac{d}{dt}\Psi^\sigma$  leads to

$$\begin{aligned} \frac{d\Psi^\sigma}{dt} &= \frac{dL}{dt}\xi^\sigma + L\frac{d\xi^\sigma}{dt} - \sum_{k=1}^N \left( \frac{d}{dt} \left( \dot{q}_k \frac{\partial L}{\partial \dot{q}_k} \right) \xi^\sigma - \dot{q}_k \frac{\partial L}{\partial \dot{q}_k} \frac{d\xi^\sigma}{dt} + \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_k} \right) \eta_k^\sigma + \frac{\partial L}{\partial \dot{q}_k} \frac{d\eta_k^\sigma}{dt} \right) - \frac{dV^\sigma}{dt} \\ &= \xi^\sigma \frac{d}{dt} \left( L - \sum_{k=1}^N \dot{q}_k \frac{\partial L}{\partial \dot{q}_k} \right) + L\frac{d\xi^\sigma}{dt} + \sum_{k=1}^N \left( \frac{\partial L}{\partial \dot{q}_k} \dot{\eta}_k^\sigma + \eta_k^\sigma \frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_k} \right) \right) - \frac{dV^\sigma}{dt} \\ &= \xi^\sigma \frac{\partial L}{\partial t} + \sum_{k=1}^N \left( \eta_k^\sigma \frac{\partial L}{\partial q_k} + \frac{\partial L}{\partial \dot{q}_k} \dot{\eta}_k^\sigma \right) + L\frac{d\xi^\sigma}{dt} - \frac{dV^\sigma}{dt} = 0. \end{aligned}$$

► first prolongation

$$\dot{\mathbf{X}}_\sigma = \xi^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}) \frac{\partial}{\partial t} + \sum_{k=1}^N \left\{ \eta_k^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}) \frac{\partial}{\partial q_k} + \dot{\eta}_k^\sigma(t, \mathbf{q}, \dot{\mathbf{q}}, \ddot{\mathbf{q}}) \frac{\partial}{\partial \dot{q}_k} \right\} \quad (43)$$

► For infinitesimal symmetry generators  $\mathbf{X}_\sigma(\epsilon_\sigma)$ , invariance is given by

$$\dot{\mathbf{X}}_\sigma(L) - \mathbf{A}(\xi^\sigma)L - \mathbf{A}(V^\sigma) = 0. \quad (44)$$

## PARTICULAR TYPES:

- Point symmetry if

$$\frac{\partial \xi^\sigma}{\partial \dot{q}_k} = \frac{\partial \eta_k^\sigma}{\partial \dot{q}_k} = 0, \quad 1 \leq k \leq N, \quad 1 \leq \sigma \leq r \quad (45)$$

- “Dynamical” symmetry if (45) does not hold.
- Variational symmetry if  $\xi(t, \mathbf{q}) = V(t, \mathbf{q}) = 0$ .
- ▶ For kinetic Lagrangians  $L = \sum_{j,k=1}^N g_{jk}(\mathbf{q}) \dot{q}_j \dot{q}_k$ , symmetry generators  $\mathbf{X}$  are Killing vectors of:

$$\bar{g} = \sum_{j,k=1}^N g_{jk}(\mathbf{q}) dq_j dq_k$$

- ▶ Related to variational symmetries.

- The necessary and sufficient condition for a Killing vector of

$$\bar{g} = \sum_{j,k=1}^N g_{jk}(\mathbf{q}) dq_j dq_k$$

to exist: there is some index  $s_0$  such that

$$\frac{\partial g_{jk}}{\partial q_{s_0}} = 0$$

for all  $1 \leq j, k \leq N$ .

- Lie algebra of Killing vectors is a subalgebra of  $\mathcal{L}_{NS}$  ( $L$  of kinetic type).
- $N$ -dimensional linearizable Lagrangian systems:  $\mathcal{L}_{NS}$  is isomorphic to the unextended Schrödinger algebra  $S(N)$ .<sup>14</sup>

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<sup>14</sup>González López A 1988 *J. Math. Physics* **29** 1097; Gorrige V. M, Leach P. G. L. 1988 *Quaest. Math.* **11** 95; Olver P J 1995 *Equivalence, Invariance and Symmetry* (Cambridge: Cambridge University Press; Wafo Soh C 2010. *Commun. Nonlinear Sci. Numer. Simulat.* **15** 139; RCS 2011 *Comm. Nonlin. Sci. Num. Sim.* **16** 3015; RCS 2012 *Comm. Nonlin. Sci. Num. Sim.* **17** 1178; Moyo S., Meleshko S. V., Ogus G. F 2013 *Comm. Nonlinear Science Num. Simulation* **18** 2972; RCS 2014 *Comm. Nonlinear Science Num. Simulation* **19** 2602; Gray R. J. 2014 *Proc. R. Soc. A* **470** 20130779; )

## INVERSE PROBLEM INVOLVING NOETHER-SYMMETRY PRESERVATION:<sup>15</sup>

- Given  $L(t, \mathbf{q}, \dot{\mathbf{q}})$  with Lie algebra  $\mathcal{L}_{NS}$ , and given  $\mathfrak{g} \subset \mathcal{L}_{NS}$ , determine perturbations

$$\widehat{L}(t, \mathbf{q}, \dot{\mathbf{q}}) = L(t, \mathbf{q}, \dot{\mathbf{q}}) + \varepsilon R(t, \mathbf{q}, \dot{\mathbf{q}}) \quad (46)$$

- subjected to the conditions
  - $\det \left( \frac{\partial^2 \widehat{L}}{\partial q^i \partial q^j} \right) = \text{cons} \neq 0$  (regularity)
  - $\mathfrak{g} \simeq \widehat{\mathcal{L}}_{NS}$  with identical generators (symmetry breaking)
  - Divergences  $V(t, \mathbf{q})$  are preserved
  - $\widehat{\psi}_X = \psi_X + \varepsilon \Psi(R(t, \mathbf{q}, \dot{\mathbf{q}}))$  for the constants of the motion
  - The system  $\widehat{L}$  is integrable in the Liouville sense
- Existence of interesting perturbations depends heavily on  $\mathfrak{g}$
- Possible solutions without Liouville integrability.

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<sup>15</sup>RCS, Guerón J 2013 *Acta Appl. Math.* **127** 105; RCS 2013 *Int. J. Geom. Meth. Phys.* **10** 130006; RCS 2014 *J. Math. Phys.* **55** 042904; RCS 2014 *Comm. Nonlin. Sci. Num. Sim.* **19** 2602; RCS 2016 *Acta Mech.* **227** 1941; RCS 2018 *Acta Mech.* **229** 211

## Generalized Ermakov-Ray-Reid systems as perturbations of the oscillator

- Damped oscillator in  $N = 2$ :  $[g_1(t), g_2(t)$  arbitrary]

$$\ddot{q}_i + g_1(t) \dot{q}_i + g_2(t) q_i = 0, \quad i = 1, 2 \quad (47)$$

- admissible Lagrangian:

$$L = \frac{1}{2} \varphi(t) (\dot{q}_1^2 + \dot{q}_2^2 - g_2(t) (q_1^2 + q_2^2)). \quad (48)$$

- Noether symmetries:  $X = \xi(t, \mathbf{q}) \frac{\partial}{\partial t} + \eta_j(t, \mathbf{q}) \frac{\partial}{\partial q_j}$  ( $1 \leq j \leq 2, k \neq j$ ):

$$\xi(t, \mathbf{q}) = \xi(t); \quad \eta_j(t, \mathbf{q}) = \frac{1}{2} \left( \dot{\xi}(t) - g_1(t) \xi(t) \right) q_j + \lambda_j^k q_k + \psi_j(t), \quad (49)$$

- $\xi(t)$  satisfies the ODE

$$\frac{d^3 \xi}{dt^3} + \left( 4g_2(t) - g_1^2(t) - 2 \frac{dg_1}{dt} \right) \frac{d\xi}{dt} + \left( 2 \frac{dg_2}{dt} - \left( \frac{dg_1}{dt} \right)^2 - g_1(t) \frac{dg_1}{dt} \right) \xi = 0. \quad (50)$$

- The more general  $\mathfrak{g} = \mathfrak{sl}(2, \mathbb{R})$ -preserving perturbation is

$$\tilde{L}_m = \frac{1}{2}\varphi(t) (\dot{q}_1^2 + \dot{q}_2^2 - g_2(t) (q_1^2 + q_2^2)) + G_{\varepsilon, m}(t, \mathbf{q}, \dot{\mathbf{q}}), \quad (51)$$

- where

$$G_\varepsilon(t, \mathbf{q}, \dot{\mathbf{q}}) = \varepsilon F\left(\frac{q_2}{q_1}, \varphi(t)(\dot{q}_2 q_1 - \dot{q}_1 q_2)\right) q_1^{-2} \varphi^{-2}(t). \quad (52)$$

- If  $g_1(t) = 0$ ,  $G$  reduces to

$$G_\varepsilon(\mathbf{q}, \dot{\mathbf{q}}) = \varepsilon F\left(\frac{q_2}{q_1}, (\dot{q}_2 q_1 - \dot{q}_1 q_2)\right) q_1^{-2}. \quad (53)$$

- In any case,  $(\tilde{\mathcal{L}}_m)_{NS} \simeq \mathfrak{sl}(2, \mathbb{R})$ .
- Auxiliary variables:  $r = q_2 q_1^{-1}$ ,  $W = q_1 \dot{q}_2 - \dot{q}_1 q_2$

$$\tilde{L} = \frac{\varphi(t)}{2} (\dot{q}_1^2 + \dot{q}_2^2 - g_2(t) (q_1^2 + q_2^2)) - \frac{\varepsilon}{q_1^2 \varphi(t)} F(r, \varphi(t) W) \quad (54)$$

► Equations of the motion:

$$\ddot{q}_1 + g_1(t) \dot{q}_1 + g_2(t) q_1 + \frac{2\varepsilon \dot{r}}{\varphi(t) q_1} \frac{\partial F}{\partial W} + \frac{\varepsilon r}{\varphi(t)^2 q_1^3} \left( \varphi(t) W \frac{\partial^2 F}{\partial r \partial W} - \frac{\partial F}{\partial r} \right) - \frac{2\varepsilon F}{\varphi(t)^2 q_1^3} + \frac{\varepsilon r (\dot{W} + g_1(t) W)}{q_1} \frac{\partial^2 F}{\partial W^2} = 0,$$

$$\ddot{q}_2 + g_1(t) \dot{q}_2 + g_2(t) q_2 + \varepsilon \frac{\frac{\partial F}{\partial r} - \varphi(t) W \frac{\partial^2 F}{\partial r \partial W}}{\varphi(t)^2 q_1^3} - \varepsilon \frac{(\dot{W} + g_1(t) W)}{q_1} \frac{\partial^2 F}{\partial W^2} = 0.$$

► “Classical” Ermakov-Ray-Reid systems:<sup>16</sup>

$$\ddot{q}_1 + \omega^2(t) q_1 - \frac{1}{q_2^3} \frac{\partial F_1}{\partial r} + \frac{W}{q_2^3} \frac{\partial^2 F_1}{\partial r \partial W} + \frac{\dot{W}}{q_2} \frac{\partial^2 F_1}{\partial W^2} = 0, \quad (55)$$

$$\ddot{q}_2 + \omega^2(t) q_2 + \frac{1}{q_2^2 q_1} \frac{\partial G_1}{\partial r} - \frac{W}{q_2^2 q_1} \frac{\partial^2 G_1}{\partial r \partial W} - \frac{\dot{W}}{q_1} \frac{\partial^2 G_1}{\partial W^2} = 0, \quad (56)$$

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<sup>16</sup>Ray J R, Reid J L 1979 *Phys. Lett. A* **71** 317; Ray J R, Reid J L 1981 *J. Math. Phys.* **22** 91; Reid JL, Ray JR 1982 *J. Phys. A: Math. Gen.* **15** 2751; Goedert J 1989 *Phys. Lett. A* **136** 391; Rogers C, Hoenselaers C, Ramgulam U 1993 *Ermakov structure in 2+1-dimensional systems. Canonical reduction* In: Ibragimov, N. H., Torrisi, M., Valenti, A. (eds.) *Modern Group Analysis: Advanced Analytical and Computational Methods in Mathematical Physics*, pp. 317 (Kluwer, Amsterdam); Govinder K S, Leach P G L 1994 *Phys. Lett. A* **186** 391; Haas F, Goedert, J 2001 *Phys. Lett. A* **279** 181.

► Helmholtz conditions:<sup>17</sup>

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}^j} \right) - \frac{\partial L}{\partial q^j} = f_{ij}(t, \mathbf{q}, \dot{\mathbf{q}}) (\ddot{q}^i - \omega^i(t, \mathbf{q}, \dot{\mathbf{q}})), \quad 1 \leq i, j \leq N \quad (57)$$

- The system (55)-(56) is Lagrangian

$$\frac{\partial^2 G_1}{\partial W^2} - r^2 \frac{\partial^2 F_1}{\partial W^2} = 0, \quad (58)$$

$$\begin{aligned} 3 \frac{\partial F_1}{\partial r} + \frac{1}{r^2} \frac{\partial G_1}{\partial r} - 3w \frac{\partial^2 F_1}{\partial r \partial W} - \frac{W}{r^2} \frac{\partial^2 G_1}{\partial r \partial W} + r \frac{\partial^2 F_1}{\partial r^2} \\ - \frac{1}{r} \frac{\partial^2 G_1}{\partial r^2} - rw \frac{\partial^3 F_1}{\partial r^2 \partial W} + \frac{W}{r} \frac{\partial^3 G_1}{\partial r^2 \partial W} = 0 \end{aligned} \quad (59)$$

► CONSEQUENCE: For  $g_1(t) = 0$  and  $g_2(t) = \omega^2(t)$ , any perturbation of the damped oscillator preserving the  $\mathfrak{sl}(2, \mathbb{R})$ -subalgebra corresponds to a generalized ERR system (55)-(56). Conversely, any Hamiltonian ERR system is a perturbation of the oscillator.

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<sup>17</sup>Inverse problem of Lagrangian mechanics. See e.g. von Helmholtz H 1887 *J. reine angew. Math.* **100** 137; Mayer A 1896 *Ber. kön. Ges. Wiss. Leipzig* **519** 1; Davis D R 1928 *Trans. Amer. Math. Soc.* **30** 710; Havas P 1956 *Bull. Am. Phys.* **1** 12; Havas P 1957 *Suppl. Nuovo Cimento* **5** 363; Vujanovic B 1971 *ATAA J.* **9** 131; Havas P 1973 *Acta Phys. Austriaca* **38**, 145; Djukić Dj S, Vujanović B 1975 *Acta Mech.* **23** 17; Engels E 1975 *Nuovo Cimento* **26B** 481; Engels E 1978 *Hadronic J.* **1** 465; Santilli R M 1978 *Foundations of Theoretical Mechanics vol. I* (Springer, New York);

## Free Lagrangian in $\mathbb{R}^N$

- Free system determined by  $L = \dot{q}^i \dot{q}_i$ .
- Noether symmetries  $\mathfrak{sl}(2, \mathbb{R}) \oplus \mathfrak{so}(N)$ :

$$X_1 = t^2 \frac{\partial}{\partial t} + \sum_{i=1}^N t q_i \frac{\partial}{\partial q_i} + \sum_{i=1}^N (q_i - t \dot{q}_i) \frac{\partial}{\partial \dot{q}_i}, \quad X_3 = \frac{\partial}{\partial t},$$

$$X_2 = -[X_1, X_3], \quad X_{i,j} = q_j \frac{\partial}{\partial q_i} - q_i \frac{\partial}{\partial q_j}, \quad 1 \leq i < j \leq N.$$

- $\mathfrak{so}(N)$ -part corresponds to Killing vectors.
- most general perturbation of  $L = \dot{q}^i \dot{q}_i$  preserving  $\mathfrak{so}(N) \oplus \mathfrak{sl}(2, \mathbb{R})$ :

$$\tilde{L} = \frac{1}{2} (\dot{q}_1^2 + \dot{q}_2^2 + \cdots + \dot{q}_N^2) - \varepsilon \frac{F \left( \left( \sum_{i=1}^N q_i^2 \right) \left( \sum_{i=1}^N \dot{q}_i^2 \right) - \left( \sum_{i=1}^N q_i \dot{q}_i \right)^2 \right)}{(q_1^2 + q_2^2 + \cdots + q_N^2)}, \quad (60)$$

- Invariance with respect to other symmetries  $\implies$  cyclic coordinates.
- Non-linearizable systems:  $\mathcal{L}'_{NS} \simeq \mathfrak{so}(m) \oplus \mathfrak{sl}(2, \mathbb{R}) < \mathfrak{so}(N) \oplus \mathfrak{sl}(2, \mathbb{R})$

► Auxiliary functions:

$$J_1 = \sum_{i=1}^m q_i \dot{q}_i, \quad J_2 = \sum_{i=1}^m q_i^2, \quad J_3 = \sum_{i=1}^m \dot{q}_i^2 \quad (61)$$

► Auxiliary variables:

$$z_j = \frac{q_{m+j}}{\sqrt{J_2}}, \quad z_{N-m+j} = \frac{\dot{q}_{m+j} J_2 - q_{m+j} J_1}{\sqrt{J_2}}, \quad 1 \leq j \leq N-m. \quad (62)$$

► Perturbations:

$$\tilde{L}_m = \frac{1}{2} (\dot{q}_1^2 + \cdots + \dot{q}_N^2) + G_{\varepsilon,m}(t, \mathbf{q}, \dot{\mathbf{q}}), \quad (63)$$

► where

$$G_{\varepsilon,m}(t, \mathbf{q}, \dot{\mathbf{q}}) = \varepsilon \frac{F(J_2 J_3 - J_1^2, z_1, \dots, z_{2N-2m})}{J_2}. \quad (64)$$

►  $\mathcal{L}'_{NS}$ -preserving potentials:

$$G_{\varepsilon,m}(t, \mathbf{q},) = \varepsilon \frac{F(z_1, \dots, z_{N-m})}{J_2}. \quad (65)$$

EXAMPLE:  $\mathcal{L}'_{NS} \simeq \mathfrak{so}(2) \oplus \mathfrak{sl}(2, \mathbb{R}) < \mathfrak{so}(3) \oplus \mathfrak{sl}(2, \mathbb{R})$  IN  $N = 3$

► Generic perturbation in cylindric reference:

$$\tilde{L} = \frac{1}{2} \left( \dot{\rho}^2 + \rho^2 \dot{\theta}^2 + \dot{\xi}^2 \right) - \frac{\varepsilon}{\rho^2} F \left( \frac{\xi}{\rho} \right), \quad (66)$$

►  $\theta$  cyclic  $\Rightarrow J = \rho^2 \dot{\theta} = \alpha = \text{cons.}$

► Reduction:

$$\tilde{L}' = \frac{1}{2} \left( \dot{\rho}^2 + \dot{\xi}^2 \right) + \frac{\alpha^2}{2\rho^2} - \frac{\varepsilon}{\rho^2} F \left( \frac{\xi}{\rho} \right). \quad (67)$$

► (67) is a perturbation of  $L_0 = \frac{1}{2} \left( \dot{\rho}^2 + \dot{\xi}^2 \right)$  [preserves  $\mathfrak{sl}(2, \mathbb{R})$ ]

► Constants of the motion:

$$H_r = \frac{1}{2} \left( \dot{\rho}^2 + \dot{\xi}^2 \right) + \frac{1}{\rho^2} \left( \varepsilon F \left( \frac{\xi}{\rho} \right) - \frac{\alpha^2}{2} \right), \quad (68)$$

$$I_1 = \frac{1}{2} \left( \xi \dot{\rho} - \dot{\xi} \rho \right)^2 + \int^{\xi \rho^{-1}} \left( z \left( 2\varepsilon F(z) - \alpha^2 \right) + \varepsilon \left( 1 + z^2 \right) \frac{dF}{dz} \right) dz. \quad (69)$$

- Question: Can perturbations alter the curvature?

## RIEMANN CURVATURE TENSOR

- $M$  Riemannian manifold. Regular metric tensor  $\bar{g} = \sum_{i,j=1}^N g_{ij}(\mathbf{q}) dq_i dq_j$ .

$$\Gamma_{ij}^h = \frac{1}{2} \sum_{k=1}^N g^{hk} \left\{ \frac{\partial g_{jk}}{\partial q_i} + \frac{\partial g_{ik}}{\partial q_j} - \frac{\partial g_{ij}}{\partial q_k} \right\}. \quad (70)$$

- Geodesic (local) equations

$$\ddot{q}_i + \sum_{j,k=1}^N \Gamma_{jk}^i \dot{q}_j \dot{q}_k = 0. \quad (71)$$

- (71) correspond to the equations of motion of the kinetic Lagrangian

$$L = \sum_{i,j=1}^N g_{ij}(\mathbf{q}) \dot{q}_i \dot{q}_j.$$

► Riemann tensor (in (0,4) tensor form)

$$R_{ijkl} = \frac{1}{2} \left( \frac{\partial^2 g_{il}}{\partial q_j \partial q_k} - \frac{\partial^2 g_{jl}}{\partial q_i \partial q_k} - \frac{\partial^2 g_{ik}}{\partial q_j \partial q_l} + \frac{\partial^2 g_{jk}}{\partial q_i \partial q_l} \right) + q_{pq} \left( \Gamma_{li}^p \Gamma_{kj}^q - \Gamma_{ki}^p \Gamma_{lj}^q \right) = g_{ih} R_{jkl}^h.$$

► Ricci (0, 2)-tensor

$$R_{ij} = \sum_{p=1}^N R_{ipj}^p = \sum_{k=1}^N \left( \frac{\partial \Gamma_{ji}^k}{\partial q_k} - \frac{\partial \Gamma_{ki}^k}{\partial q_j} \right) + \sum_{k,p=1}^N \left( \Gamma_{kp}^k \Gamma_{ji}^p - \Gamma_{jp}^k \Gamma_{ki}^p \right).$$

## CONSTANTS OF THE MOTION OF PERTURBATIONS

- Let  $J(t, \mathbf{q}, \dot{\mathbf{q}})$  be an invariant of  $L = \sum_{i,j=1}^N g_{ij}(\mathbf{q}) \dot{q}_i \dot{q}_j$ .
- $J(t, \mathbf{q}, \dot{\mathbf{q}})$  is a constant of the motion of  $\tilde{L} = L(t, \mathbf{q}, \dot{\mathbf{q}}) - U(t, \mathbf{q})$  whenever

$$\sum_{k=1}^N (-1)^k \left\{ \sum_{l=1}^N (-1)^l |\mathbf{g}_{(k,l)}| \frac{\partial U}{\partial q_l} \right\} \frac{\partial J}{\partial \dot{q}_k} = 0, \quad (72)$$

is satisfied, where  $\mathbf{g}_{(k,l)}$  denotes the minor of  $\mathbf{g}$  obtained by deletion of the  $k^{th}$  row and  $l^{th}$  column.

- Lagrange equations for  $\tilde{L}$ :

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) - \frac{\partial L}{\partial q_i} + \frac{\partial U}{\partial q_i} = 0, \quad 1 \leq i \leq N. \quad (73)$$

- For any index, the identity

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) - \frac{\partial L}{\partial q_i} = \sum_{j=1}^N \left( \frac{dg_{ij}}{dt} \dot{q}_j - \frac{1}{2} \frac{\partial g_{ij}}{\partial q_i} \dot{q}_j + g_{ij} \ddot{q}_j \right) = 0 \quad (74)$$

holds (normal form of equations of motion of  $L$ ).

- Normal equations of motion of  $\tilde{L}$

$$\tilde{\omega}_i = \omega_i + \frac{(-1)^i}{|\mathbf{g}|} \sum_{l=1}^N (-1)^l |\mathbf{g}_{(i,l)}| \frac{\partial U}{\partial q_l}, \quad 1 \leq i \leq N. \quad (75)$$

- Now  $J(t, \mathbf{q}, \dot{\mathbf{q}})$  is a constant of the motion of  $L$  implies that

$$\frac{dJ}{dt} = \frac{\partial J}{\partial t} + \sum_{i=1}^N \left( \dot{q}_i \frac{\partial J}{\partial q_i} + \omega_i \frac{\partial J}{\partial \dot{q}_i} \right) = 0. \quad (76)$$

- Supposed that  $J(t, \mathbf{q}, \dot{\mathbf{q}})$  is also an invariant of  $\tilde{L}$ :

$$\frac{dJ}{dt} = \frac{\partial J}{\partial t} + \sum_{i=1}^N \dot{q}_i \frac{\partial J}{\partial q_i} + \left( \omega_i + \frac{(-1)^i}{|\mathbf{g}|} \sum_{l=1}^N (-1)^l |\mathbf{g}_{(i,l)}| \frac{\partial U}{\partial q_l} \right) \frac{\partial J}{\partial \dot{q}_i} = 0. \quad (77)$$

- By (76), the latter reduces to the PDE for  $U(t, \mathbf{q})$ :

$$\sum_{i=1}^N \frac{(-1)^i}{|\mathbf{g}|} \left( \sum_{l=1}^N (-1)^l |\mathbf{g}_{(i,l)}| \frac{\partial U}{\partial q_l} \right) \frac{\partial J}{\partial \dot{q}_i} = 0. \quad (78)$$

- This equation is separable after the variables  $\dot{q}_i$ .

- EXAMPLE:

$$L = \frac{1}{2} (\dot{q}_1^2 + q_1^2 (\dot{q}_2^2 + \sin^2 q_2 \dot{q}_3^2))$$

in spherical coordinates.

- Quadratic invariant  $J = q_1^4 (\dot{q}_2^2 + \sin^2 q_2 \dot{q}_3^2)$ .
- Only nonzero entries of the matrix  $\mathbf{g}$  are  $g_{11} = 1$ ,  $g_{22} = q_1^2$ ,  $g_{33} = q_1^2 \sin^2 q_2$
- Only nonvanishing minors are

$$|\mathbf{g}_{(1,1)}| = q_1^4 \sin^2 q_2, \quad |\mathbf{g}_{(2,2)}| = q_1^2 \sin^2 q_2, \quad |\mathbf{g}_{(3,3)}| = q_1^2.$$

- Resulting PDE for  $L - U(t, \mathbf{q})$ :

$$2q_1^6 \sin^2 q_2 \left( \dot{q}_2 \frac{\partial U}{\partial q_2} + \dot{q}_3 \frac{\partial U}{\partial q_3} \right) = 0. \quad (79)$$

- Separable equation with solution potentials  $U(\mathbf{q}) = V(t, q_1)$ .

## Systems containing dissipative terms

- Lagrangian  $L = \sum_{i,j=1}^N g_{ij}(\mathbf{q}) \dot{q}_i \dot{q}_j$ .
- $J(t, \mathbf{q}, \dot{\mathbf{q}})$  invariant of  $L$  homogeneous of order  $\alpha$  in the velocities  $\dot{\mathbf{q}}$ .
- $\tilde{J} = f(t)^\alpha J(t, \mathbf{q}, \dot{\mathbf{q}})$  is an invariant of  $\tilde{L} = f(t)L$ .
- Normal form  $\ddot{q}_i = \tilde{\omega}_i(t, \mathbf{q}, \dot{\mathbf{q}})$  of the EM of  $\tilde{L}$

$$\tilde{\omega}_i(t, \mathbf{q}, \dot{\mathbf{q}}) = \omega_i(t, \mathbf{q}, \dot{\mathbf{q}}) - \frac{d}{dt}(\ln f(t)) \dot{q}_i, \quad 1 \leq i \leq N. \quad (80)$$

- Total differential of  $\tilde{J}$

$$\frac{d\tilde{J}}{dt} = \dots = f(t)^\alpha \frac{dJ}{dt} + \dot{f}(t) f(t)^{\alpha-1} \left( \alpha J - \sum_{i=1}^N \dot{q}_i \frac{\partial J}{\partial \dot{q}_i} \right) = 0. \quad (81)$$

- First term vanishes (invariance of  $J$ ).
- Second term is zero by the Euler theorem on homogeneous functions.

- Extension to Lagrangians with a potential are possible by considering invariants as a sum of homogeneous functions in the  $\dot{\mathbf{q}}$  of decreasing degree, and modifying it according to these orders.
- $\mathbf{X}$  Killing vector of

$$\overline{ds^2} = \sum_{i,j=1}^N g_{ij}(\mathbf{q}) dq_i dq_j$$

associated to the kinetic Lagrangian  $L = \sum_{i,j=1}^N g_{ij}(\mathbf{q}) \dot{q}_i \dot{q}_j$ .

- $\mathbf{X}$  is a Noether point symmetry of  $\tilde{L} = f(t)L$  for any non-zero function  $f(t)$ . As  $\mathbf{X}$  is a Killing vector, we have that  $\xi(t, \mathbf{q}) = V(t, \mathbf{q}) = 0$  and  $\frac{\partial \eta_i}{\partial t} = 0$  for  $1 \leq i \leq n$  in the symmetry condition (44). Hence

$$\dot{\mathbf{X}}(\tilde{L}) = \dot{\mathbf{X}}(f(t)L) = f(t) \sum_{i=1}^N \left( \eta_i \frac{\partial L}{\partial q_i} + \dot{\eta}_i \frac{\partial L}{\partial \dot{q}_i} \right) = f(t) \dot{\mathbf{X}}(L) = 0.$$

## PLANE PERTURBATIONS

- Most obvious starting case

$$L_n = \frac{q_1^n}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2) \quad (82)$$

- $n = 0$ :  $L_0$  free Lagrangian
- $n \neq 0$ : can be seen as a position-dependent mass system.
- Symmetry generators  $\mathbf{X} = \xi(t, \mathbf{q}) \frac{\partial}{\partial t} + \eta^1(t, \mathbf{q}) \frac{\partial}{\partial q_1} + \eta^2(t, \mathbf{q}) \frac{\partial}{\partial q_2}$

$$\begin{aligned} \xi(t, \mathbf{q}) &= b_1 t^2 + b_2 t + b_3, \\ \eta^1(t, \mathbf{q}) &= \frac{(2b_1 t + b_2)}{n+2} q_1 + q_1^{-\frac{n}{2}} \left\{ (C_{11}t + C_{21}) \sin \left( \frac{n+2}{2} q_2 \right) + (C_{12}t + C_{22}) \cos \left( \frac{n+2}{2} q_2 \right) \right\}, \\ \eta^2(t, \mathbf{q}) &= -q_1^{-\frac{n+2}{2}} \left\{ (C_{12}t + C_{22}) \sin \left( \frac{n+2}{2} q_2 \right) + (C_{11}t + C_{21}) \cos \left( \frac{n+2}{2} q_2 \right) \right\} + b_4. \end{aligned}$$

- Gauge term

$$V(t, \mathbf{q}) = \frac{2}{n+2} q_1^{\frac{n+2}{2}} \left( C_{11} \sin \left( \frac{n+2}{2} q_2 \right) + C_{12} \cos \left( \frac{n+2}{2} q_2 \right) \right) + \frac{2b_1}{n+2} q_1^{n+2},$$

- $\mathcal{L}_{NS}$  isomorphic  $S(2)$
- Levi part of isomorphic to  $\mathfrak{sl}(2, \mathbb{R})$

$$X_1 = t^2 \frac{\partial}{\partial t} + \frac{2t q_1}{n+2} \frac{\partial}{\partial q_1}, \quad X_2 = t \frac{\partial}{\partial t} + \frac{q_1}{n+2} \frac{\partial}{\partial q_1}, \quad X_3 = \frac{\partial}{\partial t}. \quad (83)$$

- Hamiltonian  $H$  constant of the motion.
- Symmetries corresponding to  $C_{21}, C_{22}, b_4$  variational.
- Most general term quadratic in  $\dot{\mathbf{q}}$

$$R(t, \mathbf{q}, \dot{\mathbf{q}}) = q_1^{(n+2)} S(q_2) \dot{q}_2^2 + T(q_2) \dot{q}_2 + U(q_2) q_1^{-(n+2)}, \quad (84)$$

- $T(q_2) = 0$ , as a total time derivative.
- Regularity ensured if

$$1 + 2S(q_2) \neq 0.$$

- Riemann curvature obviously zero.

- For generic functions  $S(q_2)$  and  $U(q_2) \neq 0$

$$L_{n,\varepsilon} = \frac{q_1^n}{2} (\dot{q}_1^2 + q_1^2 (1 + 2S(q_2)) \dot{q}_2^2) + U(q_2) q_1^{-(n+2)}$$

is a solution to the inverse problem with  $\mathfrak{g} = \mathfrak{sl}(2, \mathbb{R})$ .

- Constants of the motion: Hamiltonian and

$$J_2 = \frac{q_1^{2n+4}}{2} (1 + 2S(q_2)) \dot{q}_2^2 + U(q_2). \quad (85)$$

## INTRODUCTION OF TIME-DEPENDENT FACTOR

$$\tilde{L}_n = \varphi(t) \frac{q_1^n}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2), \quad n \neq -2. \quad (86)$$

- System still linearizable with  $\dim \mathcal{L}_{NS} = 8$ .

## ► Symmetries

$$\begin{aligned}\xi(t, \mathbf{q}) &= G(t), \\ \eta^1(t, \mathbf{q}) &= q_1^{-\frac{n}{2}} \left\{ \sin\left(\frac{n+2}{2}q_2\right) \left( C_1 + C_2 \int \frac{dt}{\varphi(t)} \right) + \cos\left(\frac{n+2}{2}q_2\right) \left( C_3 + C_4 \int \frac{dt}{\varphi(t)} \right) \right\} + \frac{(\varphi(t)\dot{G}(t) - \dot{\varphi}(t)G(t))}{(n+2)\varphi(t)} q_1, \\ \eta^2(t, \mathbf{q}) &= q_1^{-\frac{n+2}{2}} \left\{ \cos\left(\frac{n+2}{2}q_2\right) \left( C_1 + C_2 \int \frac{dt}{\varphi(t)} \right) - \sin\left(\frac{n+2}{2}q_2\right) \left( C_3 + C_4 \int \frac{dt}{\varphi(t)} \right) \right\} + C_5\end{aligned}$$

## ► Gauge term

$$\begin{aligned}V(t, \mathbf{q}) &= \frac{2}{n+2} q_1^{\frac{n+2}{2}} \left( C_2 \sin\left(\frac{n+2}{2}q_2\right) + C_4 \cos\left(\frac{n+2}{2}q_2\right) \right) + \frac{2}{n+2} q_1^{n+2} \times \\ &\quad \left( \dot{\varphi}(t)^2 G(t) + \varphi(t)^2 \ddot{G}(t) - \dot{\varphi}(t) \dot{G}(t) \varphi(t) - \varphi(t) G(t) \ddot{\varphi}(t) \right).\end{aligned}$$

## ► Constraint on function $G(t)$ :

$$\ddot{z}(t) - \frac{\left( 2\ddot{\varphi}(t)\varphi(t) - \dot{\varphi}(t)^2 \right)}{4\varphi(t)^2} z(t) = 0. \quad (87)$$

## ► Symmetries for which $\xi(t, \mathbf{q}) = V(t, \mathbf{q}) = 0$ are preserved.

- Lagrangians solving the inverse problem

$$\tilde{L}_{n,\varepsilon} = \varphi(t) \frac{q_1^n}{2} (\dot{q}_1^2 + q_1^2 (1 + 2\varepsilon S(q_2)) \dot{q}_2^2) - \varepsilon \frac{U(q_2)}{\varphi(t) q_1^{n+2}}. \quad (88)$$

- $H$  not a constant of the motion, but  $\tilde{J}_1 = \varphi(t) H$ .

- Second invariant

$$\tilde{J}_2 = \varphi(t)^2 \frac{q_1^{2n+4}}{2} (1 + 2\varepsilon S(q_2)) \dot{q}_2^2 + \varepsilon U(q_2) \quad (89)$$

- $\tilde{J}_2 \neq \varphi(t) J_2$ , as

$$\varphi(t) (L_n + \varepsilon R(t, \mathbf{q}, \dot{\mathbf{q}})) - \tilde{L}_{n,\varepsilon} = \varepsilon \frac{U(q_2)}{q_1^{n+2}} \frac{(\varphi^2(t) - 1)}{\varphi(t)} \neq 0,$$

- Symmetry-preserving perturbation and multiplication by time functions  $\varphi(t)$  not interchangeable.

## A PDM-VERSION OF ERMAKOV-RAY-REID SYSTEMS

- Generalized Ermakov-Ray-Reid systems (ERR) admitting a Hamiltonian arise as symmetry-preserving perturbations of the oscillator  $\ddot{\mathbf{x}} + \omega^2(t) \mathbf{x} = 0$  with time-dependent frequency.
- In polar coordinates  $q_1 = \rho$ ,  $q_2 = \theta$ , we obtain  $L_0 = \frac{1}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2 - \omega^2(t) q_1^2)$ .
- Generalization to ERR-systems with pdm. Let  $n \neq -2$

$$L_n = \frac{\varphi(t) q_1^n}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2 - \omega^2(t) q_1^2). \quad (90)$$

- Inverse problem with  $\mathfrak{sl}(2, \mathbb{R})$ -preservation

$$L_{n,\varepsilon} = \frac{\varphi(t) q_1^n}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2 - \omega^2(t) q_1^2) + \varepsilon \varphi(t)^{-1} q_1^{-(n+2)} \Phi(q_2, \varphi(t) q_1^{n+2} \dot{q}_2), \quad (91)$$

- with

$$\Phi(q_2, \varphi(t) q_1^{n+2} \dot{q}_2)$$

arbitrary function of its arguments.

- New variable  $w = \varphi(t) q_1^{n+2} \dot{q}_2$ : two independent constants of the motion are

$$\begin{aligned} J_1 &= \varphi(t) \left( \dot{q}_1 \frac{\partial L_{n,\varepsilon}}{\partial \dot{q}_1} + \dot{q}_2 \frac{\partial L_{n,\varepsilon}}{\partial \dot{q}_2} - L_{n,\varepsilon} \right) - \frac{1}{2} q_1^{n+2} \varphi^2(t) \omega^2(t) \\ &\quad - \int^{q_1^{n+2}} (\varphi(\zeta) \omega(\zeta))^2 d\zeta, \end{aligned} \quad (92)$$

$$J_2 = \varphi^2(t) q_1^{2n+4} \dot{q}_2^2 + 2\varepsilon \varphi(t) q_1^{n+2} \dot{q}_2 \frac{\partial \Phi}{\partial w} - 2\varepsilon \Phi(q_2, w) \quad (93)$$

- The choice  $\varphi(t) = 1$  provides a pdm-version of the generalized ERR systems.
- We recover the two invariants of the classical ERR systems as a limit:

$$\begin{aligned} \lim_{n \rightarrow 0} J_1 &= \frac{1}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2) + \varepsilon \dot{q}_2 \frac{\partial \Phi}{\partial w} - \varepsilon q_1^{-2} \Phi(q_2, w) + \frac{1}{2} \int^{q_1^2} \omega^2(\zeta) d\zeta, \\ \lim_{n \rightarrow 0} J_2 &= q_1^4 \dot{q}_2^2 + 2\varepsilon q_1^2 \dot{q}_2 \frac{\partial \Phi}{\partial w} - 2\varepsilon \Phi(q_2, w). \end{aligned} \quad (94)$$

## A FUNCTIONAL APPROACH

- $\theta(t), \rho(t)$  arbitrary functions,  $\{u(t), v(t)\}$  independent solutions of

$$\ddot{z}(t) + \rho(t)\dot{z}(t) + \theta(t)z(t) = 0. \quad (95)$$

- $\xi(t) = C_1u(t)^2 + C_2u(t)v(t) + C_3v(t)^2$  determines the general solution of

$$\ddot{\xi} + 3\dot{\rho}\ddot{\xi} + (\dot{\rho} + 2\rho^2 + 4\theta)\dot{\xi} + (4\rho\theta + 2\dot{\theta})\xi = 0. \quad (96)$$

- Consider vector fields

$$\mathbf{X} = \xi(t) \frac{\partial}{\partial t} + \frac{1}{2} \dot{\xi}(t) q_i \frac{\partial}{\partial q_i},$$

- $\mathbf{X}_i$  ( $1 \leq i \leq 3$ ) is associated to  $C_j = \delta_i^j$ .

- Commutators .

$$\begin{aligned} [\mathbf{X}_1, \mathbf{X}_2] &= W(u, v)\mathbf{X}_1, \quad [\mathbf{X}_1, \mathbf{X}_3] = 2W(u, v)\mathbf{X}_2, \\ [\mathbf{X}_2, \mathbf{X}_3] &= W(u, v)\mathbf{X}_3, \end{aligned} \quad (97)$$

$W(u, v) = uv - \dot{u}\dot{v}$  Wronskian.

- If  $W(u, v) = \text{cons.}$  we recover  $\mathfrak{sl}(2, \mathbb{R})$ .

- Define

$$V(\mathbf{q}) = \frac{1}{4} \ddot{\xi}(t) (q_1^2 + q_2^2),$$

- $X_i$  are Noether point symmetries of  $L(t, \mathbf{q}, \dot{\mathbf{q}})$  if

$$\begin{aligned} \frac{1}{2} \frac{\partial L}{\partial \dot{q}_1} (\ddot{\xi} q_1 - \dot{\xi} \dot{q}_1) + \frac{1}{2} \frac{\partial L}{\partial \dot{q}_2} (\ddot{\xi} q_2 - \dot{\xi} \dot{q}_2) + \xi \frac{\partial L}{\partial t} + \frac{\dot{\xi}}{2} \left( q_1 \frac{\partial L}{\partial q_1} + q_2 \frac{\partial L}{\partial q_2} \right) \\ - \frac{\ddot{\xi}}{4} (q_1^2 + q_2^2) - \frac{1}{2} \ddot{\xi} \dot{q}_1 q_1 - \frac{1}{2} \ddot{\xi} \dot{q}_2 q_2 + \dot{\xi} L = 0 \end{aligned}$$

- General solution

$$\begin{aligned} L(t, \mathbf{q}, \dot{\mathbf{q}}) = & \frac{(\xi \ddot{\xi} - \dot{\xi}^2)}{4\xi^2} (q_1^2 + q_2^2) + \frac{\dot{\xi} (q_1 \dot{q}_1 + q_2 \dot{q}_2)}{2\xi} \\ & + \frac{1}{\xi} \Psi \left( \frac{q_1}{\sqrt{\xi}}, \frac{q_2}{\sqrt{\xi}}, \frac{\dot{\xi} q_1 - 2\dot{q}_1 \xi}{\sqrt{\xi}}, \frac{\dot{\xi} q_2 - 2\dot{q}_2 \xi}{\sqrt{\xi}} \right), \end{aligned}$$

- Contains the class of oscillatory systems

$$L = \frac{1}{2} (\dot{q}_1^2 + \dot{q}_2^2) + \frac{2\xi(\ddot{t})\xi(t) - 1}{\xi(t)^2} (q_1^2 + q_2^2). \quad (98)$$

## PERTURBED PDM-SYSTEMS IN $N = 3$

- Starting point

$$L_n = \frac{q_1^n}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2 + q_1^2 \sin^2(q_2) \dot{q}_3^2) \quad (99)$$

- associated to the metric tensor

$$\overline{ds^2} = \frac{q_1^n}{2} (dq_1^2 + q_1^2 dq_2^2 + q_1^2 \sin^2(q_2) dq_3^2). \quad (100)$$

- nonzero components of  $R_{ijkl}$ , up to symmetries

$$R_{2323} = -\frac{n(4+n)}{2} q_1^{n+2} \sin^2(q_2) \quad (101)$$

- the Riemann tensor vanishes only if  $n = 0$  or  $n = -4$ .
- Lagrangian  $L_n$  is linearizable only for  $n = 0, -4$ .
- Number of Noether point symmetries is 12.

►  $\mathcal{L}_{NS}$  Lie algebra of Noether point symmetries.

1. If  $n = 0, -4$ , then  $\mathcal{L}_{NS}$  is isomorphic to the unextended Schrödinger algebra  $S(3)$ .
2. If  $n = -2$ , then  $\mathcal{L}_{NS}$  is isomorphic to the direct sum  $\mathfrak{so}(3) \oplus \mathfrak{h}_1$ , where  $\mathfrak{h}_1$  denotes the three dimensional Heisenberg algebra.
3. If  $n \neq -4, -2, 0$ , then  $\mathcal{L}_{NS} \simeq \mathfrak{sl}(2, \mathbb{R}) \oplus \mathfrak{so}(3)$  and coincides with the Levi subalgebra of  $S(3)$ .

► The value  $n = -2$  must be analyzed separately, Rotational symmetry (Killing vectors) preserved, remaining Noether point symmetries

$$\begin{aligned}\xi(t, \mathbf{q}) &= C_1, & \eta^1(t, \mathbf{q}) &= C_2 t q_1 + C_3 q_1, \\ \eta^2(t, \mathbf{q}) &= C_4 \sin q_3 + C_5 \cos q_3, & \eta^3(t, \mathbf{q}) &= C_4 \cos q_3 \cot q_2 - C_5 \sin q_3 \cot q_2 + C_6.\end{aligned}$$

►  $C_1, C_2$  and  $C_3$  generate the Heisenberg algebra  $\mathfrak{h}_1$ .

- The most general perturbation  $\tilde{L}_n = L_n + \varepsilon R(t, \mathbf{q}, \dot{\mathbf{q}})$  preserving  $\mathfrak{sl}(2, \mathbb{R})$  is

$$R(t, \mathbf{q}, \dot{\mathbf{q}}) = \Phi(q_2, q_3, \dot{q}_2 q_1^{n+2}, \dot{q}_3 q_1^{n+2}) q_1^{-(n+2)}. \quad (102)$$

- Imposing the symmetry condition leads to

$$\begin{aligned} \frac{\partial R}{\partial t} &= 0 \\ \frac{t^2(n+2)}{2} \frac{\partial R}{\partial t} + t q_1 \frac{\partial R}{\partial q_1} + (q_1 - t \dot{q}_1 (n+1)) \frac{\partial R}{\partial \dot{q}_1} - t(n+2) + \\ &\left( \dot{q}_2 \frac{\partial R}{\partial \dot{q}_2} + \dot{q}_3 \frac{\partial R}{\partial \dot{q}_3} \right) + t(n+2) R(t, \mathbf{q}, \dot{\mathbf{q}}) = 0. \end{aligned}$$

- simplified to

$$q_1 \frac{\partial R}{\partial q_1} - (n+2) \left( \dot{q}_2 \frac{\partial R}{\partial \dot{q}_2} + \dot{q}_3 \frac{\partial R}{\partial \dot{q}_3} \right) + (n+2) R(\mathbf{q}, \dot{\mathbf{q}}) = 0, \quad q_1 \frac{\partial R}{\partial \dot{q}_1} = 0. \quad (103)$$

- For  $\tilde{L}_{n,\varepsilon}$  with  $\varphi_4(q_2, q_3) = 0$ . If

$$\frac{\partial \varphi_i}{\partial q_3} \neq 0$$

no symmetries corresponding to Killing vectors.

- To ensure enough independent invariants, larger subgroup required.

EXAMPLE:

$$\tilde{L}_{n,\varepsilon} = \frac{q_1^n}{2} \left\{ \dot{q}_1^2 + q_1^2 \left( (1 + \varepsilon \varphi_1(q_2)) \dot{q}_2^2 + 2\varepsilon \varphi_2(q_2) \dot{q}_2 \dot{q}_3 + (\sin^2(q_2) + \varepsilon \varphi_3(q_2)) \dot{q}_3^2 \right) \right\} + \frac{\varepsilon \varphi_4(q_2)}{q_1^{n+2}}.$$

- Admitted invariants

$$J_1 = \frac{q_1^n}{2} \left\{ \dot{q}_1^2 + q_1^2 \left( (1 + 2\varepsilon \varphi_1(q_2)) \dot{q}_2^2 + 4\varepsilon \varphi_2(q_2) \dot{q}_2 \dot{q}_3 + (\sin^2(q_2) + 2\varepsilon \varphi_3(q_2)) \dot{q}_3^2 \right) \right\} - \frac{\varepsilon \varphi_4(q_2)}{q_1^{n+2}},$$

$$J_2 = q_1^{n+2} (2\varepsilon \varphi_2(q_2) \dot{q}_2 + (\sin^2 q_2 + 2\varepsilon \varphi_2(q_2) \dot{q}_3)),$$

$$J_3 = q_1^{2n+4} ((1 + 2\varepsilon \varphi_1(q_2)) \dot{q}_2^2 + 4\varepsilon \varphi_2(q_2) \dot{q}_2 \dot{q}_3 + (\sin^2(q_2) + 2\varepsilon \varphi_3(q_2)) \dot{q}_3^2) - 2\varepsilon \varphi_4(q_2).$$

- Higher orders in  $\dot{q}$  can be obtained by dynamical symmetries.

## PERTURBATIONS ASSOCIATED TO FLAT METRICS

- For any  $n \neq -2$ ,

$$L_n = \frac{q_1^n}{2} (\dot{q}_1^2 + q_1^2 \dot{q}_2^2 + q_1^2 \sin^2(q_2) \dot{q}_3^2) \quad (104)$$

admits perturbations with  $\mathcal{L}_{NS} \simeq \mathfrak{sl}(2, \mathbb{R}) \oplus \mathbb{R}$  and flat metric:

$$\tilde{R}_{ijkl} = 0, \quad 1 \leq i, j, k, l \leq 3.$$

- As special solution among perturbations we find

$$\tilde{L}_{n,\varepsilon} = \frac{q_1^n}{2} \left\{ \dot{q}_1^2 + q_1^2 \left( \dot{q}_2^2 + 2\varepsilon \varphi_2(q_2) \dot{q}_2 \dot{q}_3 + \sin^2(q_2) \dot{q}_3^2 \right) \right\}, \quad (105)$$

- Curvature tensor

$$\begin{aligned} \tilde{R}_{2323} &= \frac{n(4+n)\sin^4 q_2 + \varepsilon^2 \left( n(4+n)\cos(2q_2) - (2+n)^2 \right) \varphi_2^2(q_2) + \varepsilon^4 (2+n)^4 \varphi_2^4(q_2)}{q_1^{n+2} (\varepsilon \varphi_2^2(q_2) - \sin^2 q_2)^2} \\ &\quad + \frac{2\varepsilon^2 \sin(2q_2) \varphi_2(q_2) \frac{d\varphi_2}{dq_2}}{q_1^{n+2} (\varepsilon \varphi_2^2(q_2) - \sin^2 q_2)^2}. \end{aligned} \quad (106)$$

- Parameterized nonlinear first order differential equation for  $\varphi_2(q_2)$ :<sup>18</sup>

$$\frac{n(4+n)\sin^4 q_2}{\varphi_2(q_2)} + \varepsilon^2 \left( n(4+n)\cos(2q_2) - (2+n)^2 \right) \varphi_2(q_2) + \varepsilon^4 (2+n)^4 \varphi_2^3(q_2) + 2\varepsilon^2 \sin(2q_2) \frac{d\varphi_2}{dq_2} = 0.$$

- Solution

$$\varphi_2(q_2) = \varepsilon^{-1} \frac{\sqrt{C_1 + n(4+n)\cos^2 q_2}}{\sqrt{(2+n)^2 \cot^2 q_2 + C_1 \csc^2 q_2}}, \quad C_1 \in \mathbb{R}. \quad (107)$$

- Stronger assertion: If

$$R_{ij} = 0 \Rightarrow R_{ijkl} = 0.$$

- Solutions of perturbations correspond to geodesic lines of flat metrics.

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<sup>18</sup>Appell M P 1889 *Journ. de Math.* **5** 361.

## PDM-SYSTEMS WITH DISSIPATIVE TERMS

- Lagrangian of Kepler system

$$L = \frac{1}{2} \left( \dot{q}_1^2 + q_1^2 \dot{q}_2^2 + q_1^2 \sin^2(q_2) \dot{q}_3^2 + \frac{\alpha}{q_1} \right), \quad (108)$$

- Obvious generalization

$$L_n = \frac{q_1^n}{2} \left( \dot{q}_1^2 + q_1^2 \dot{q}_2^2 + q_1^2 \sin^2(q_2) \dot{q}_3^2 + \frac{\alpha}{q_1} \right). \quad (109)$$

- Restricting to Noether point symmetries

1. If  $n \neq 1$ , then  $\dim \mathcal{L}_{NS} = 4$  with  $\mathcal{L}_{NS} \simeq \mathbb{R} \oplus \mathfrak{so}(3)$ .
2. If  $n = 1$ , then  $\dim \mathcal{L}_{NS} = 6$  with  $\mathcal{L}_{NS} \simeq \mathfrak{sl}(2, \mathbb{R}) \oplus \mathfrak{so}(3)$ .

- Are there superintegrable solutions?

- Special value  $n = 1$

$$\tilde{L}_1 = \frac{q_1}{2} \left( \dot{q}_1^2 + q_1^2 (1 + 2\gamma) (\dot{q}_2^2 + \sin^2(q_2) \dot{q}_3^2) + \frac{\alpha}{q_1} \right) + \frac{\beta}{q_1^3}.$$

- $\mathfrak{so}(3)$ -preserving perturbation term

$$R(t, \mathbf{q}, \dot{\mathbf{q}}) = \Phi(t, q_1, \dot{q}_1, \dot{q}_2^2 + \sin^2 q_2 \dot{q}_3^2). \quad (110)$$

- $\mathcal{L}_{NS}$ -preserving perturbation term

$$\Phi(q_1, w) = \frac{1}{q_1^3} \Psi(q_1^6 w) = \frac{1}{q_1^3} \Psi(q_1^6 (\dot{q}_2^2 + \sin^2 q_2 \dot{q}_3^2)). \quad (111)$$

- Quadratic in  $\dot{\mathbf{q}}$

$$R(t, \mathbf{q}, \dot{\mathbf{q}}) = C_1 q_1^3 (\dot{q}_2^2 + \sin^2 q_2 \dot{q}_3^2) + \frac{\beta}{q_1^3}, \quad (112)$$

- Constants of the motion:

$$\begin{aligned} J_1 &= \frac{q_1}{2} \left\{ \dot{q}_1^2 + q_1^2 (1 + 2\gamma) (\dot{q}_2^2 + \sin^2(q_2) \dot{q}_3^2) - \frac{\alpha}{q_1} \right\} - \frac{\beta}{q_1^3} \\ J_2 &= q_1^3 \sin^2 q_2 \dot{q}_3, \\ J_3 &= q_1^6 (\dot{q}_2^2 + \sin^2 q_2 \dot{q}_3^2), \\ J_4 &= 2 \sin q_3 \dot{q}_2 + \cos q_3 \sin(2q_2) \dot{q}_3, \end{aligned}$$

- No solutions being maximally superintegrable.
- Other generalizations

$$\tilde{L}_1 = \Psi(\mathbf{q}, \dot{\mathbf{q}}) \left( \dot{q}_1^2 + q_1^2 (1 + 2\gamma) (\dot{q}_2^2 + \sin^2(q_2) \dot{q}_3^2) + \frac{\alpha}{q_1} \right) + U(\mathbf{q})$$

possible, however not being maximally superintegrable.

## APPROACH VIA REALIZATIONS OF VF

- Generic  $\mathfrak{sl}(2, \mathbb{R})$  realization

$$X_1 = t^2 \frac{\partial}{\partial t} + 2\alpha q_1^k t \frac{\partial}{\partial q_1}, \quad X_2 = t \frac{\partial}{\partial t} + \alpha q_1^k \frac{\partial}{\partial q_1}, \quad X_3 = \frac{\partial}{\partial t}, \quad (113)$$

- NPS of Lagrangians

$$L = \frac{1}{2} \exp \left( \frac{q_1^{1-k}}{\alpha(1-k)} \right) \left( q_1^{-2k} \varphi_1(\mathbf{q}') \dot{q}_1^2 + 2\alpha q_1^{-k} \sum_{j=1}^N \frac{\partial \varphi_1}{\partial q_j} \dot{q}_1 \dot{q}_j + \sum_{j,k=2}^N \varphi_{ij}(\mathbf{q}') \dot{q}_j \dot{q}_k \right) - \exp \left( \frac{-q_1^{1-k}}{\alpha(1-k)} \right) U(\mathbf{q}'),$$

$$\mathbf{q}' = (q_2, \dots, q_N), \quad \varphi_1(\mathbf{q}'), \varphi_{ij}(\mathbf{q}'), U(\mathbf{q}')$$

- Gauge term

$$V(t, \mathbf{q}) = 2\alpha^2 \varphi_1(\mathbf{q}') \exp \left( \frac{q_1^{1-k}}{\alpha(1-k)} \right). \quad (114)$$

- $N = 2, U(q_2) = 0$  metric

$$\overline{ds^2} = \frac{1}{2} \exp\left(\frac{q_1^{1-k}}{\alpha(1-k)}\right) \left( q_1^{-2k} \varphi_1(q_2) \dot{q}_1^2 + 2\alpha q_1^{-k} \frac{d\varphi_1}{dq_2} \dot{q}_1 \dot{q}_2 + \varphi_{22}(q_2) \dot{q}_2^2 \right),$$

$$\varphi_{12}(q_2) \varphi_1(q_2) - 4 \left( \frac{d\varphi_1}{dq_2} \right)^2 \neq 0.$$

- Riemann tensor

$$R_{1212} = -\frac{1}{8} \exp\left(\frac{q_1^{1-k}}{\alpha(1-k)}\right) \frac{\left(2\frac{d^2\varphi_1}{dq_2^2}\varphi_1(q_2) - \frac{d\varphi_1}{dq_2}^2\right) \varphi_{12}(q_2) - \varphi_1(q_2) \frac{d\varphi_1}{dq_2} \frac{d\varphi_{12}}{dq_2}}{2\alpha^2 \frac{d\varphi_1}{dq_2}^2 - \varphi_1(q_2) \varphi_{12}(q_2)},$$

- Vanishes only if

$$\varphi_{12}(q_2) = \lambda \left( \frac{d\varphi_1}{dq_2} \right)^2 / \varphi_1(q_2).$$

## AN $n$ -DIMENSIONAL EXTRAPOLATION

- Metric

$$ds^2 = \frac{1}{(q^n)^2} (dq^1 \otimes dx^1 + \cdots + dq^n \otimes dq^n) \quad (115)$$

- Conformal group  $SO(1, n)$  as isometry group.
- $\mathcal{L}_{PS} \simeq \mathfrak{so}(1, n) \oplus \mathfrak{r}_2$  [ $\mathfrak{r}_2$  the 2-dimensional affine Lie algebra]
- The vector field  $\mathbf{Y} = t \frac{\partial}{\partial t}$  leads to

$$-\frac{1}{2q_n^2} \sum_{k=1}^n \left( \dot{q}_k^2 + \dot{q}_k \frac{\partial V}{\partial q_k} \right) - \frac{\partial V}{\partial t}, \quad (116)$$

no gauge term  $V(t, \mathbf{q})$  possible.

- $\mathcal{L}_{NS} \simeq \mathfrak{so}(1, n) \oplus \mathbb{R}$
- System ( $L = L_0 - U(t, \mathbf{q})$ )

$$\ddot{q}_k = \frac{2\dot{q}_k \dot{q}_n}{q_n} - \frac{q_n^2}{2} \frac{\partial U}{\partial q_k}, \quad 1 \leq k \leq n-1; \quad \ddot{q}_n = \frac{\dot{q}_n^2 - \dot{q}_1^2 - \cdots - \dot{q}_{n-1}^2}{q_n} - \frac{q_n^2}{2} \frac{\partial U}{\partial q_n}$$

- $\mathcal{L}_{NS} \subset \mathfrak{so}(1, n) \oplus \mathbb{R}$  is a proper subalgebra.

## Some final remarks. Possible outlines.

- Systematization of inverse problems for “pure” dynamical symmetries.
  - Symmetry formulation using (involutive) distributions.
  - Harrison-Estabrook formalism.
  - Connection with classification of Lie algebras as VF on manifolds.
  - Extension to Lagrangian densities and PDE systems.
  - Additional criteria for linearization criteria by **non-point** symmetries.
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